

# NILC results for PICO

Mathieu Remazeilles

3 June 2021

# To-do list

## ☐ Model 90.91

- **Baseline (21-800 GHz):**
  - $r = 0$  (10 realizations of CMB and noise)
  - $r = 0.003$  (10 realizations of CMB and noise)
- **w/o LF (43-800 GHz):**
  - $r = 0$  (10 realizations of CMB and noise)
- **w/o HF (21-462 GHz):**
  - $r = 0$  (10 realizations of CMB and noise)

## ☐ Model 90.92

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## ☐ Model MKD

# To-do list

## ☐ Model 90.91

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✓ done!

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- **w/o LF (43-800 GHz):**

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✓ done!

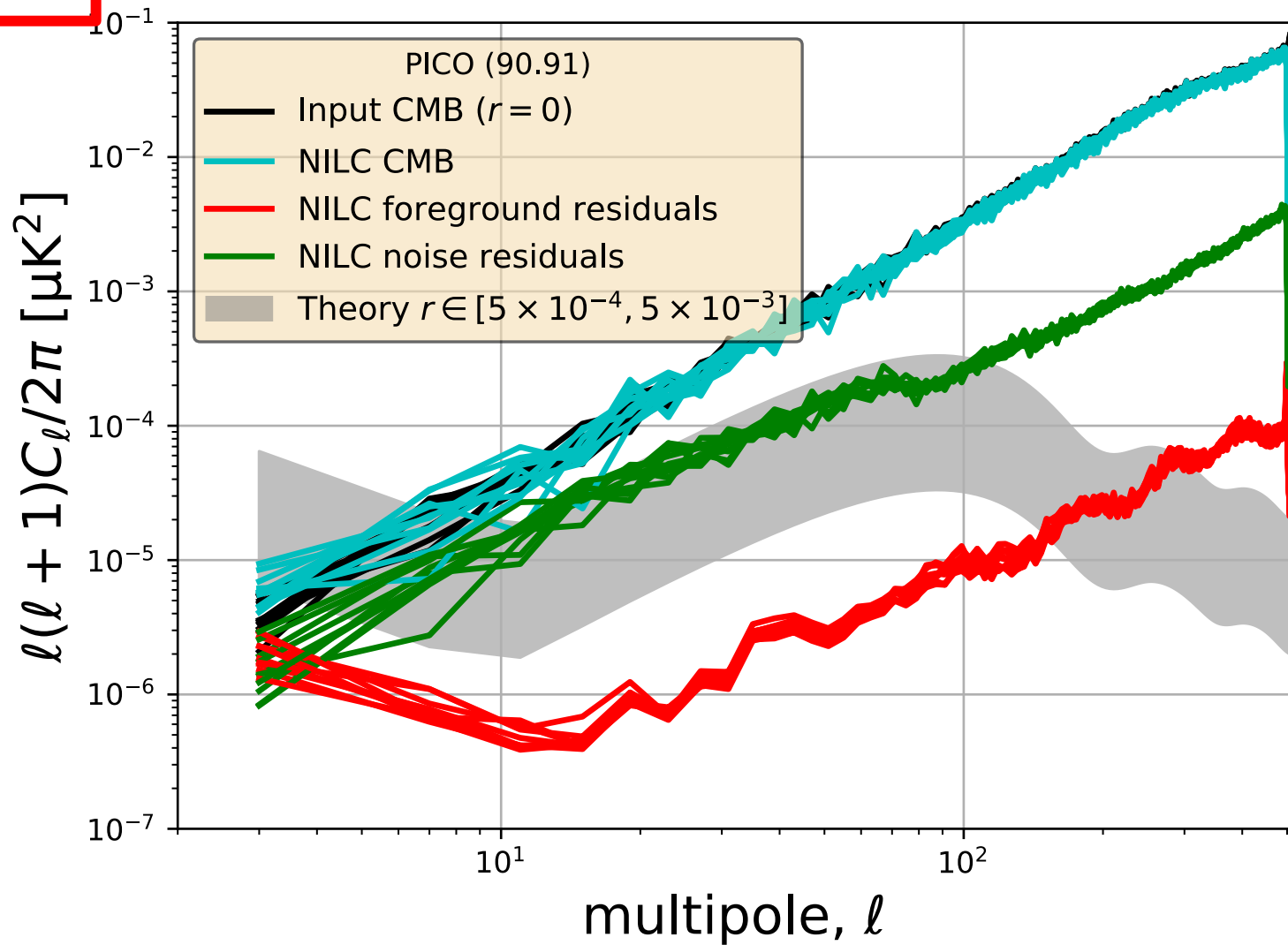
## ☐ Model MKD

$$r = 0$$

90.91 & 90.92

Baseline  
21-800 GHz

90.91,  $r = 0$   
NILC



10 realizations

MASTER

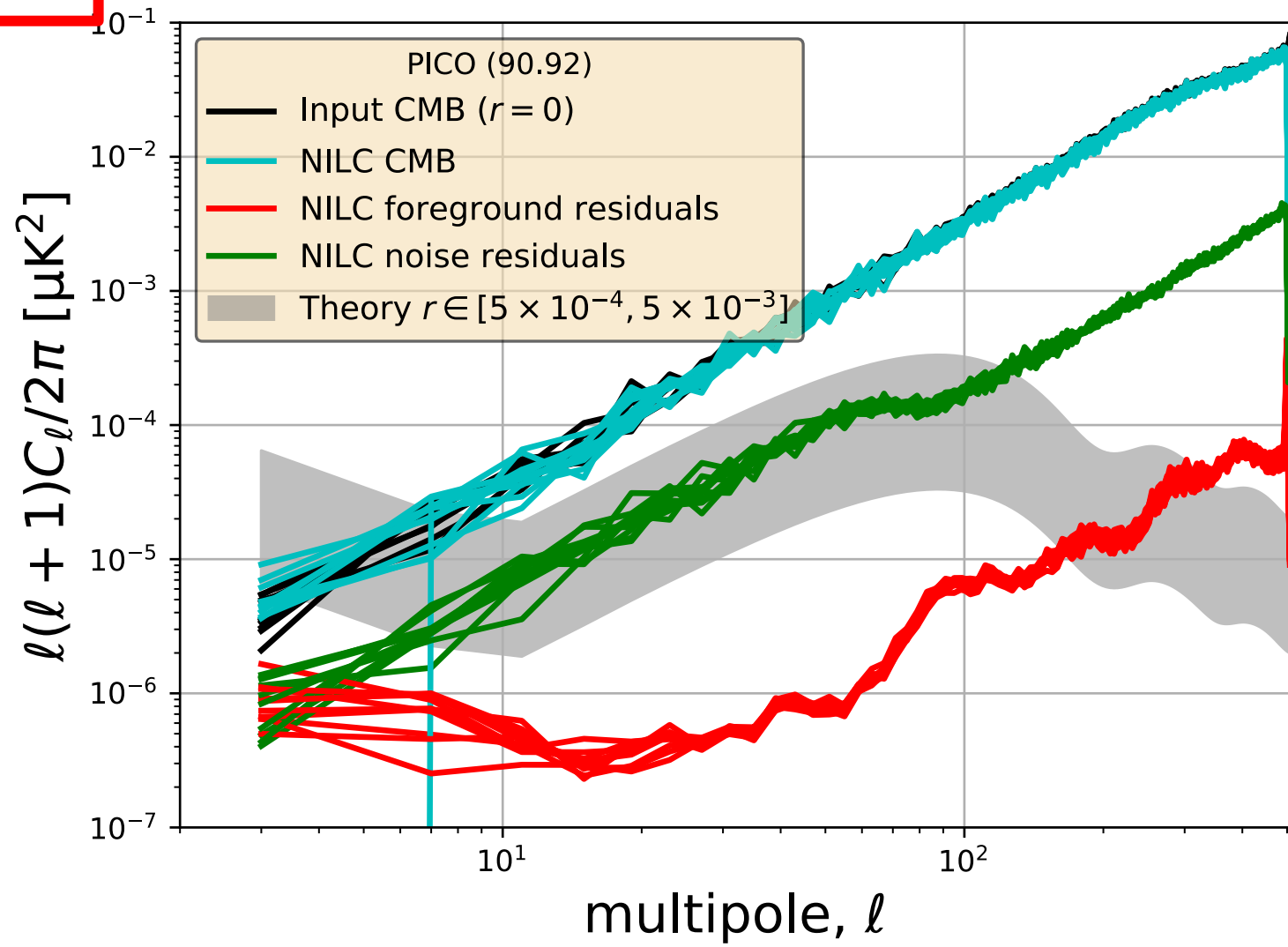
$f_{\text{sky}} = 50\%$

Binning:  $\Delta\ell = 4$

Baseline  
21-800 GHz

90.92,  $r = 0$   
NILC

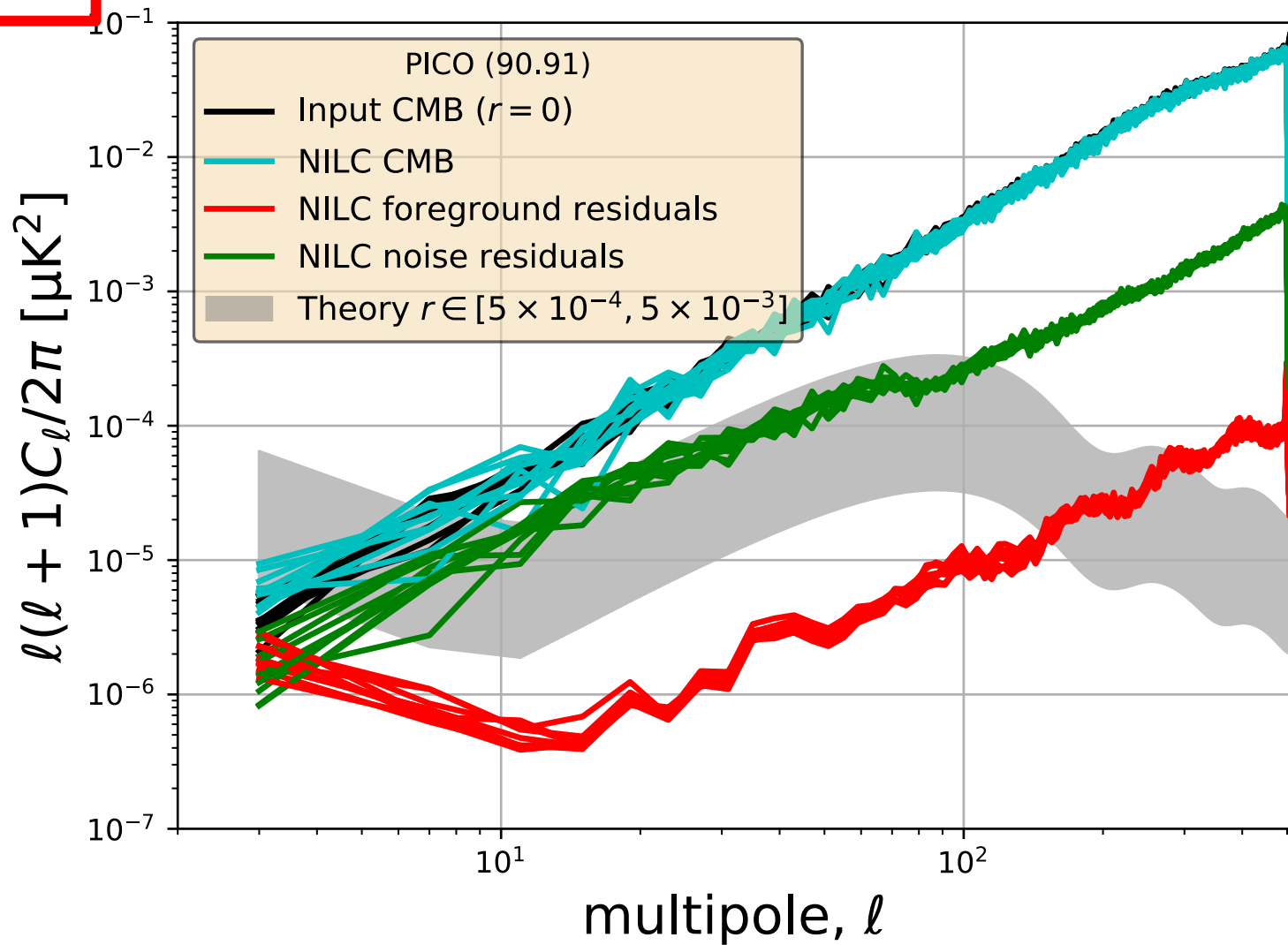
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10 realizations

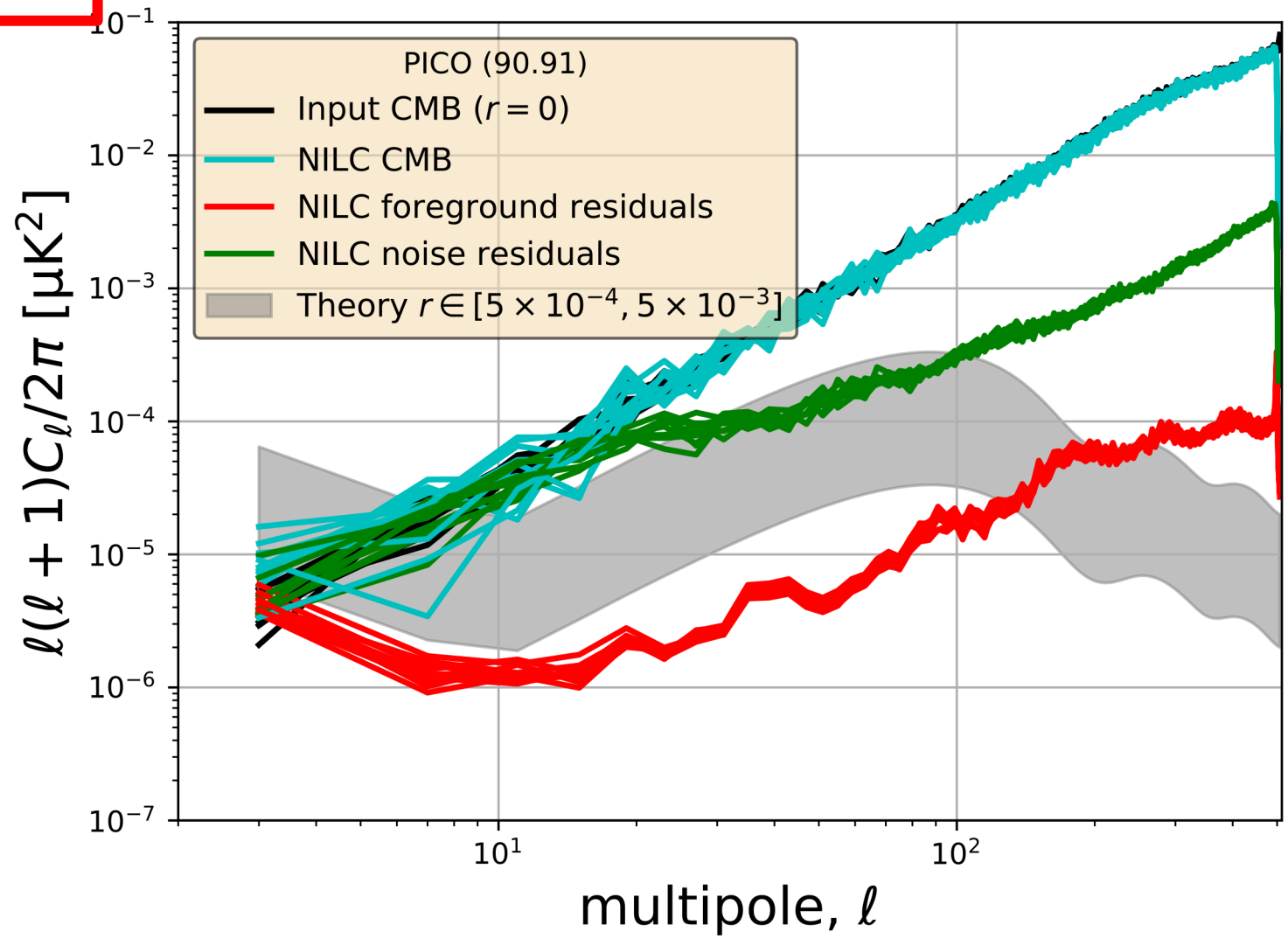
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w/o LF  
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NILC

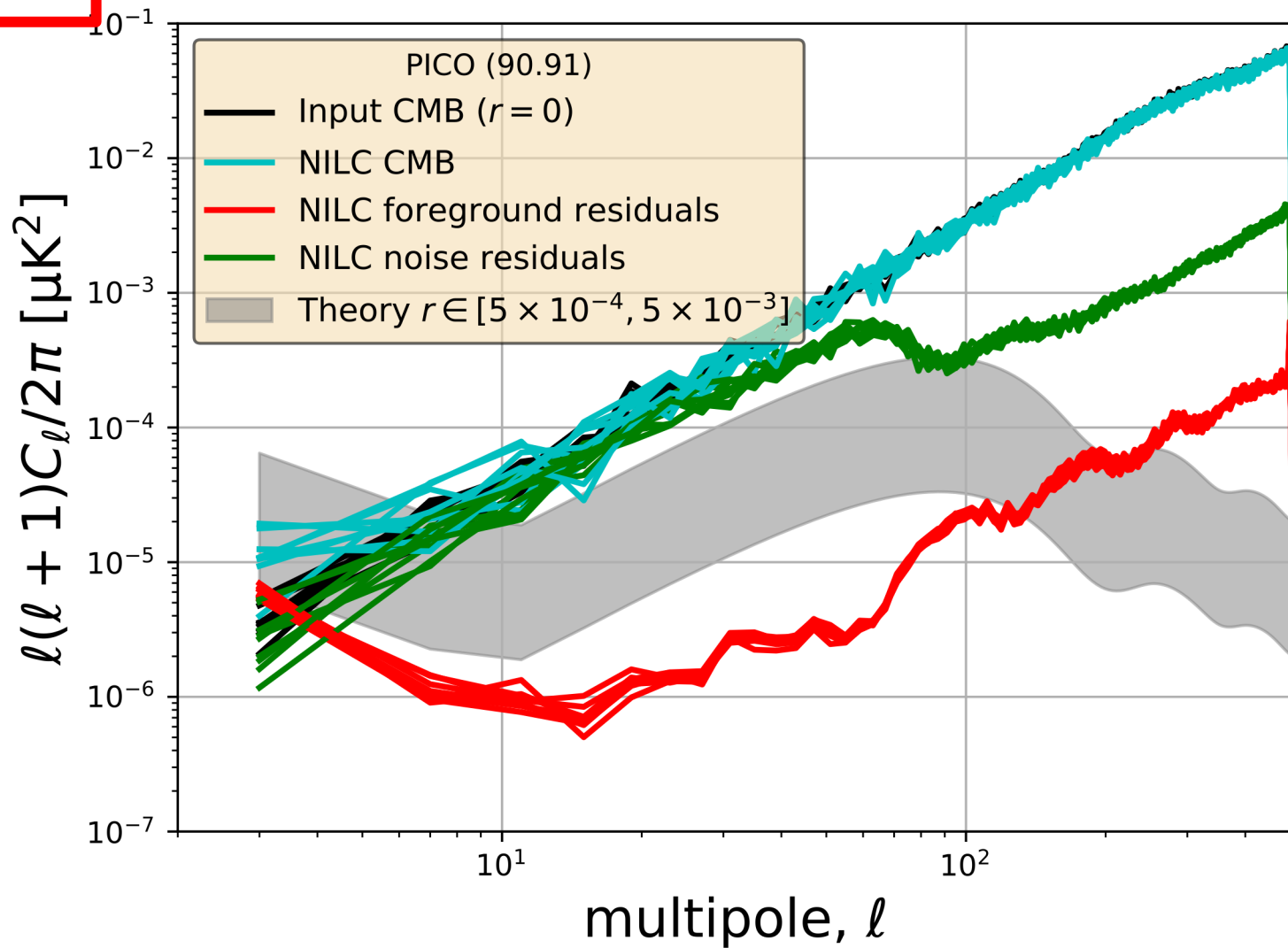


10 realizations

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 $f_{\text{sky}} = 50\%$   
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w/o HF  
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NILC



10 realizations

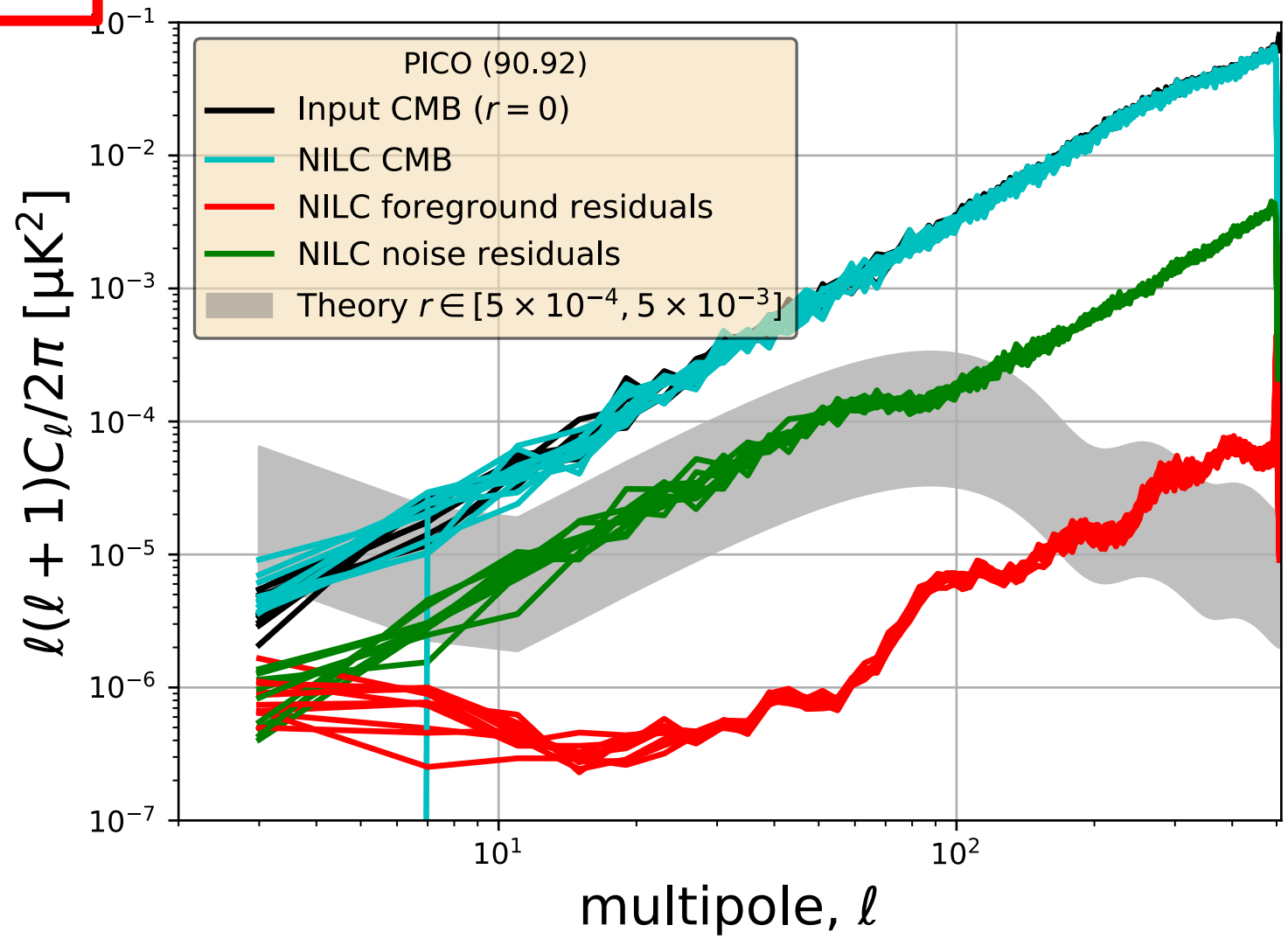
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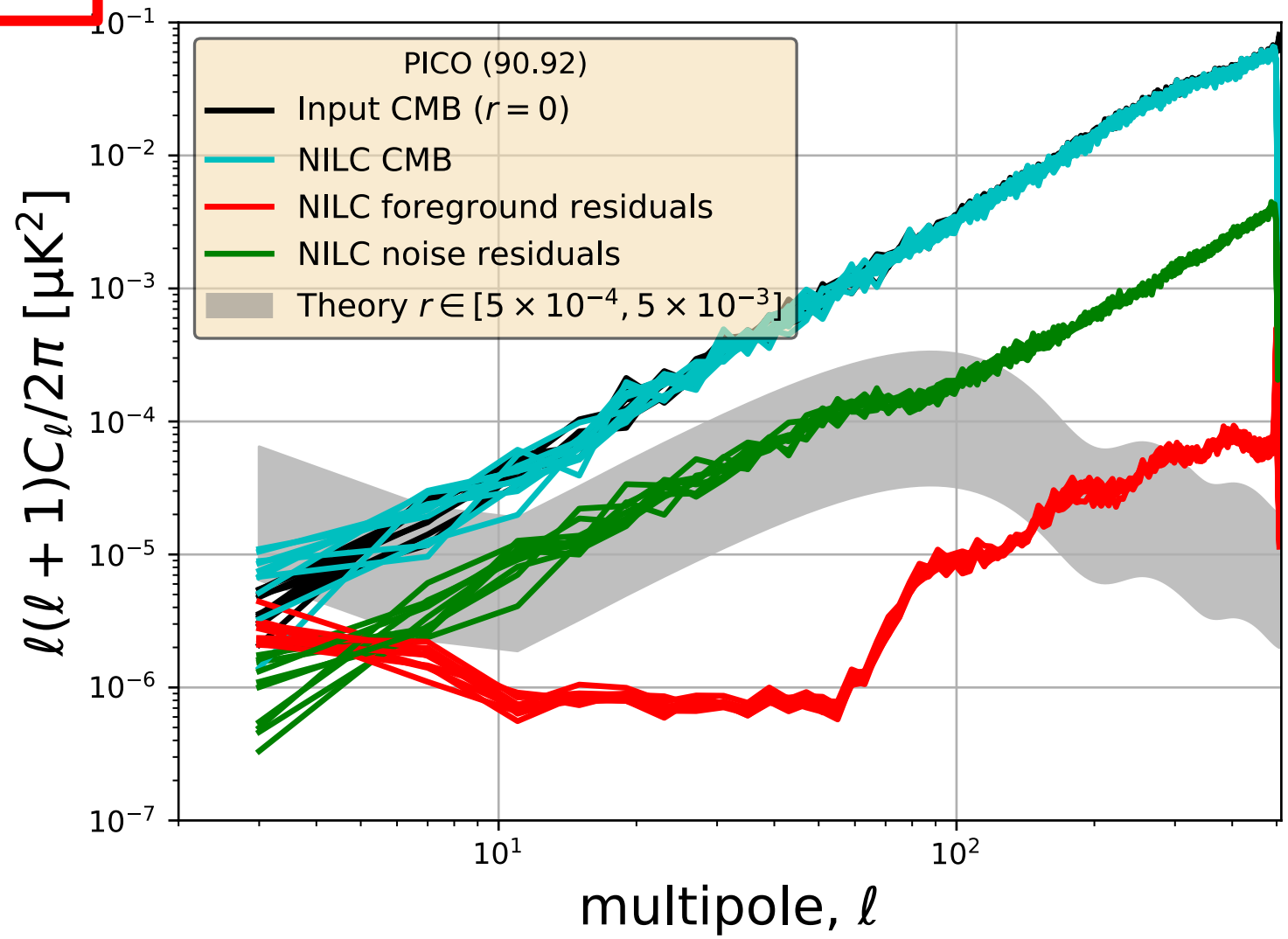


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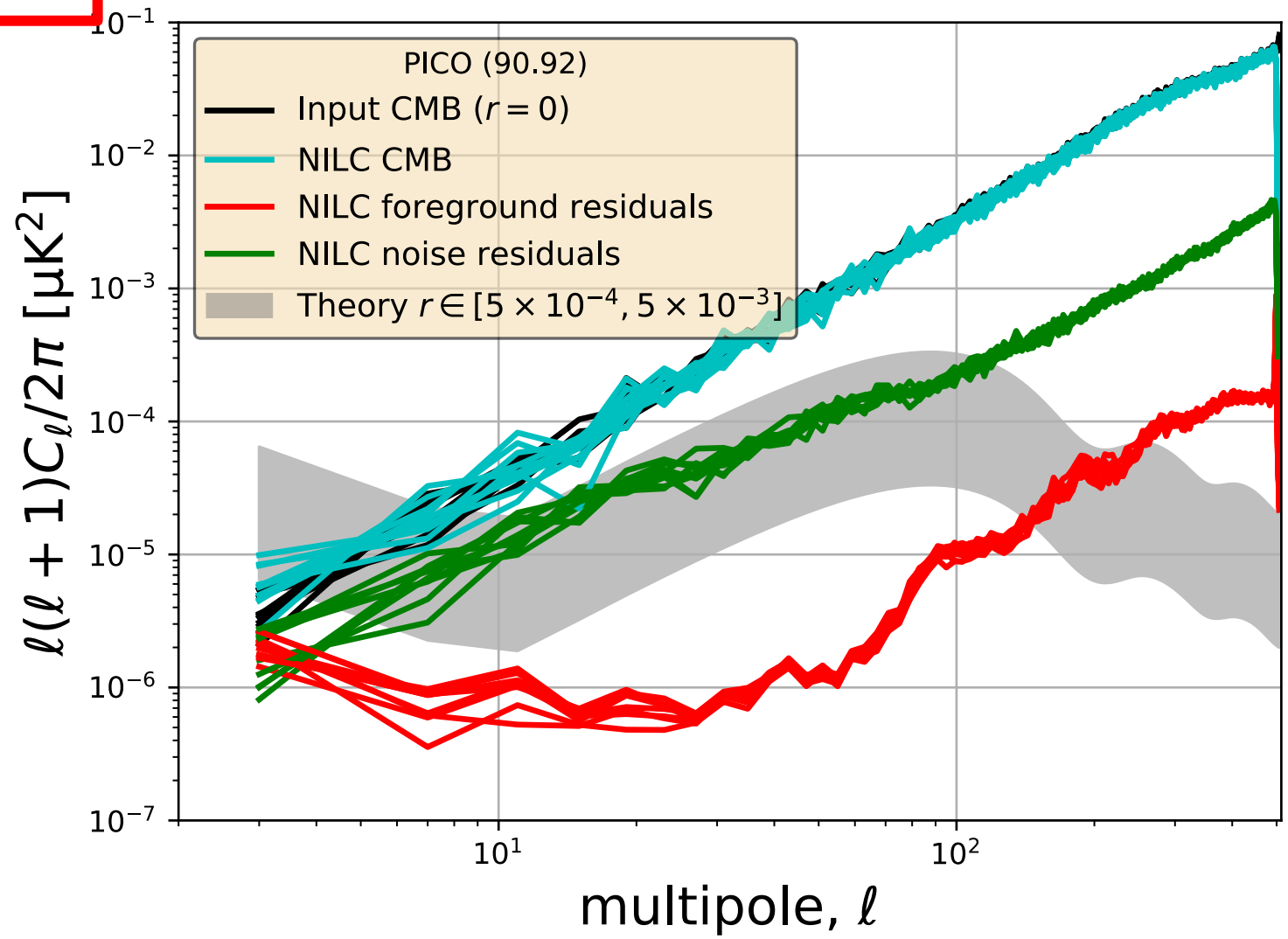


10 realizations

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21-462 GHz

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10 realizations

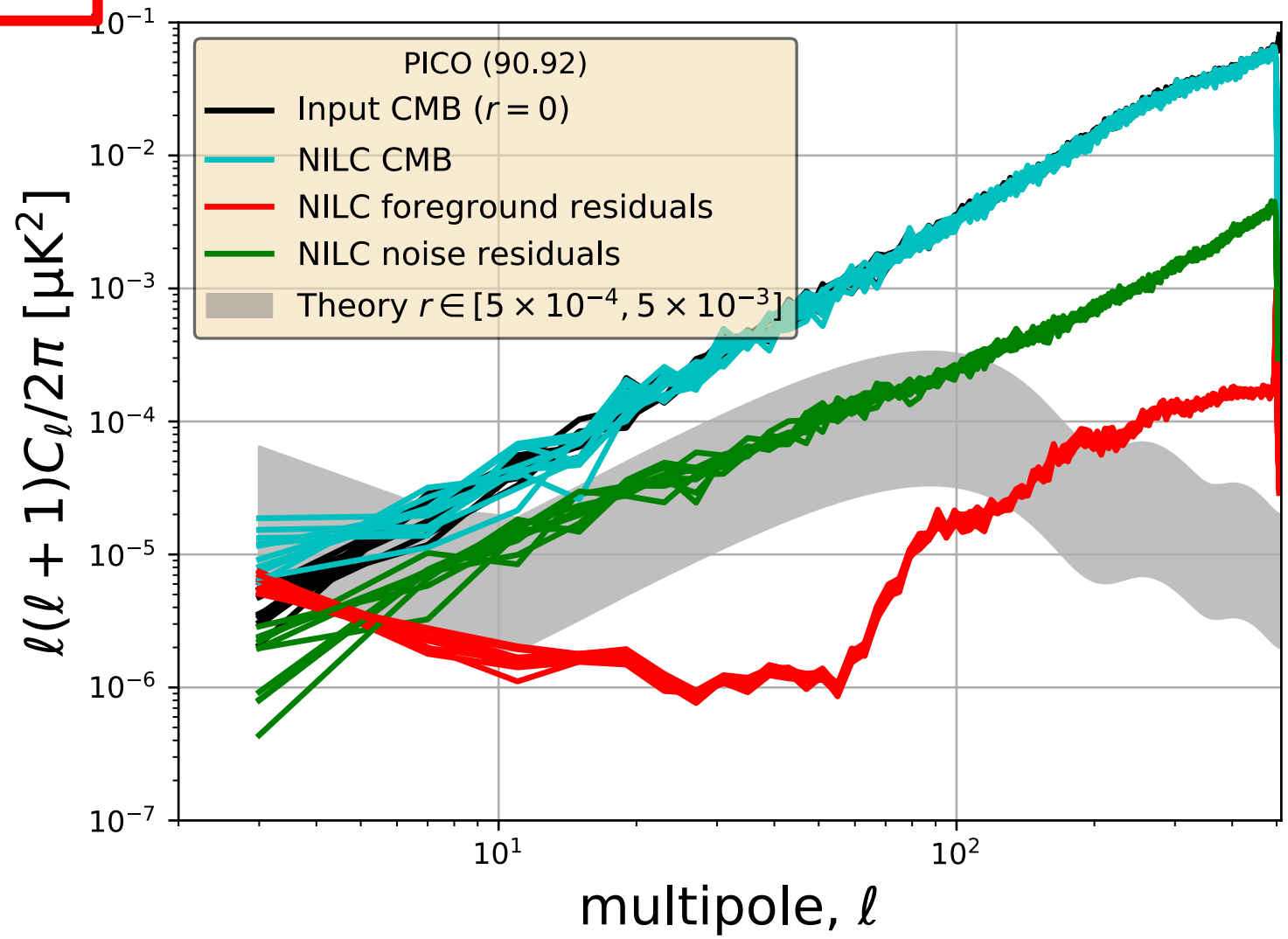
MASTER

$f_{\text{sky}} = 50\%$

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Descope  
43-462 GHz

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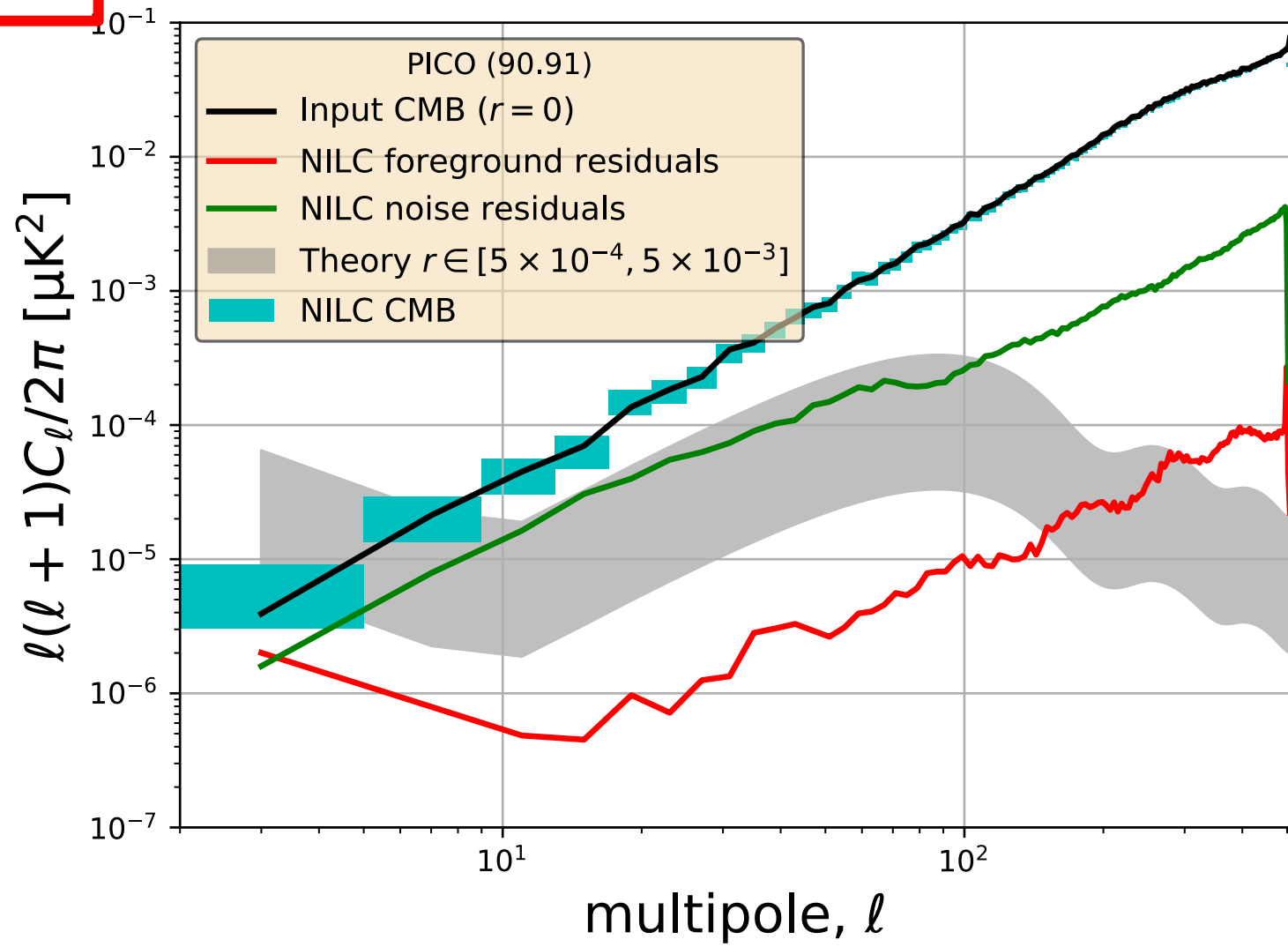
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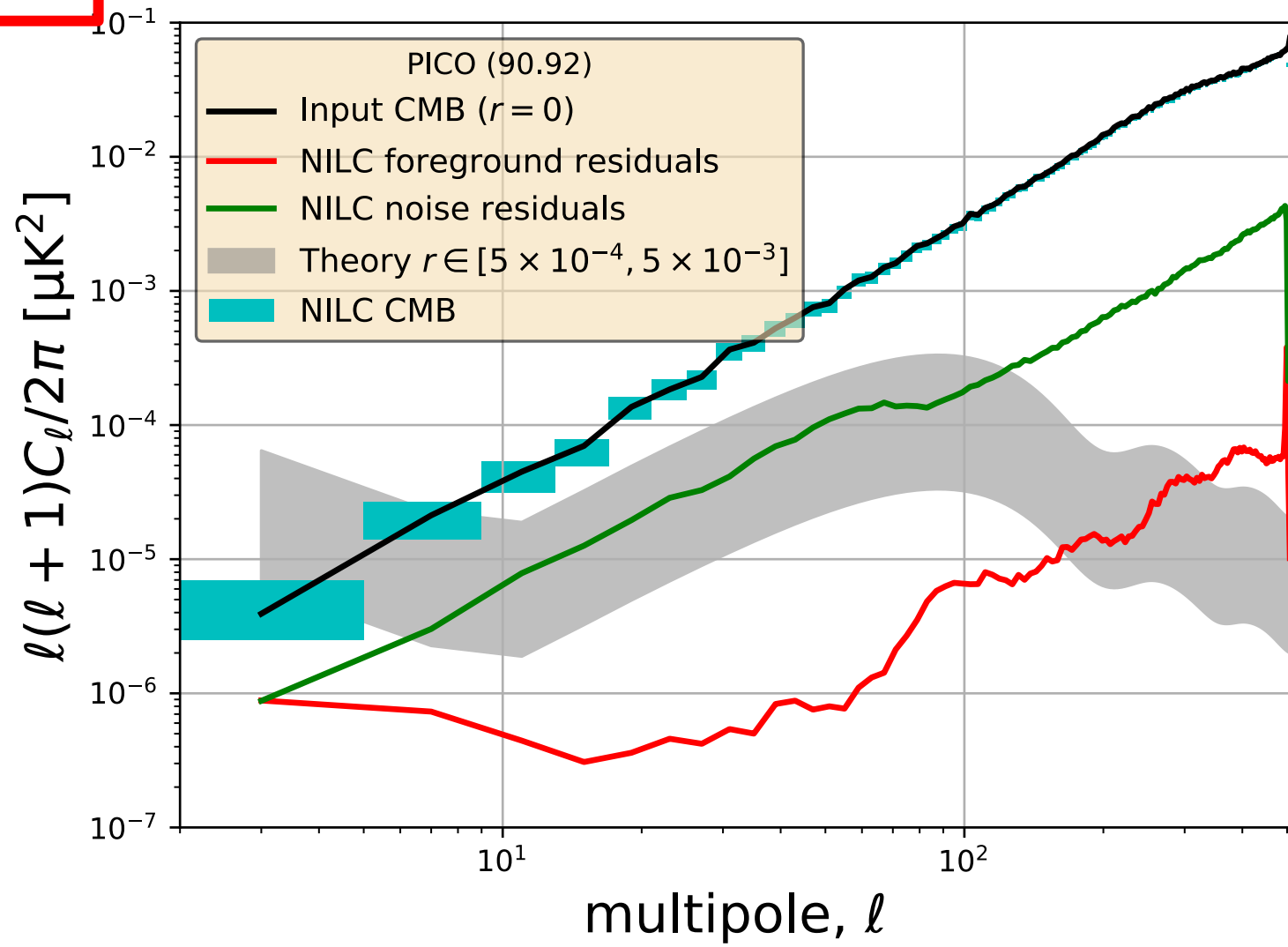


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NILC

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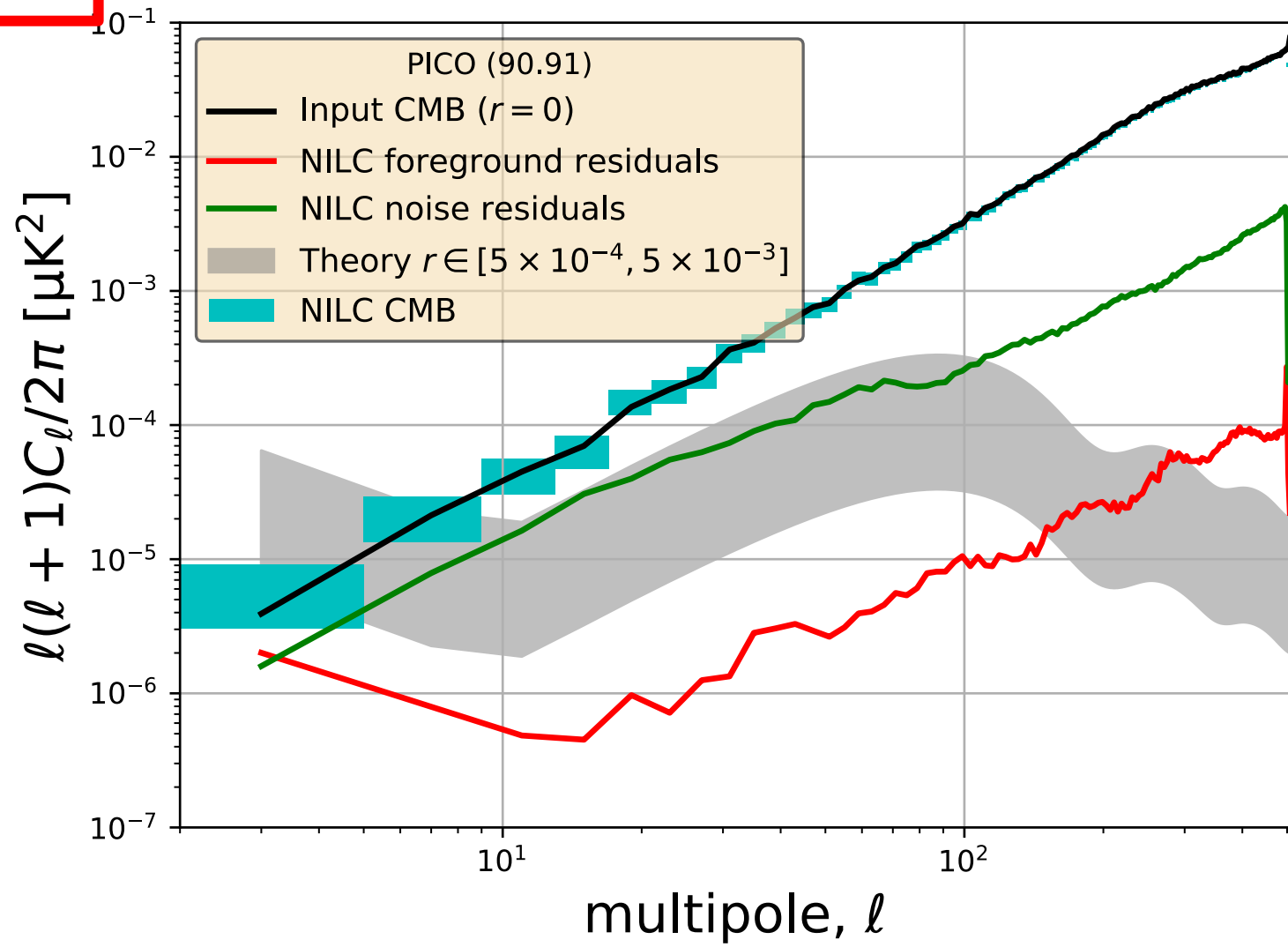


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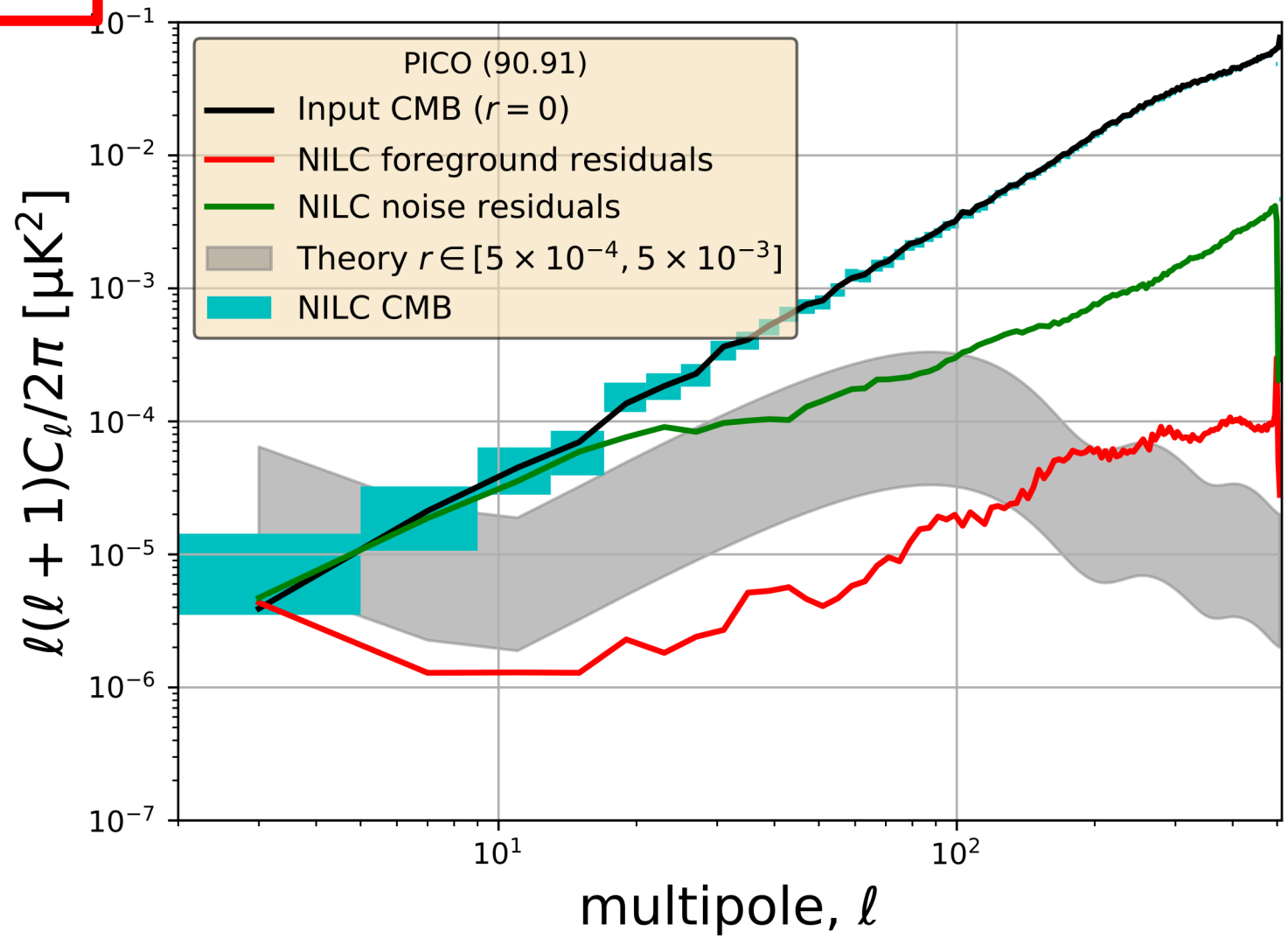
10 realizations



MASTER  
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w/o LF  
43-800 GHz

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NILC



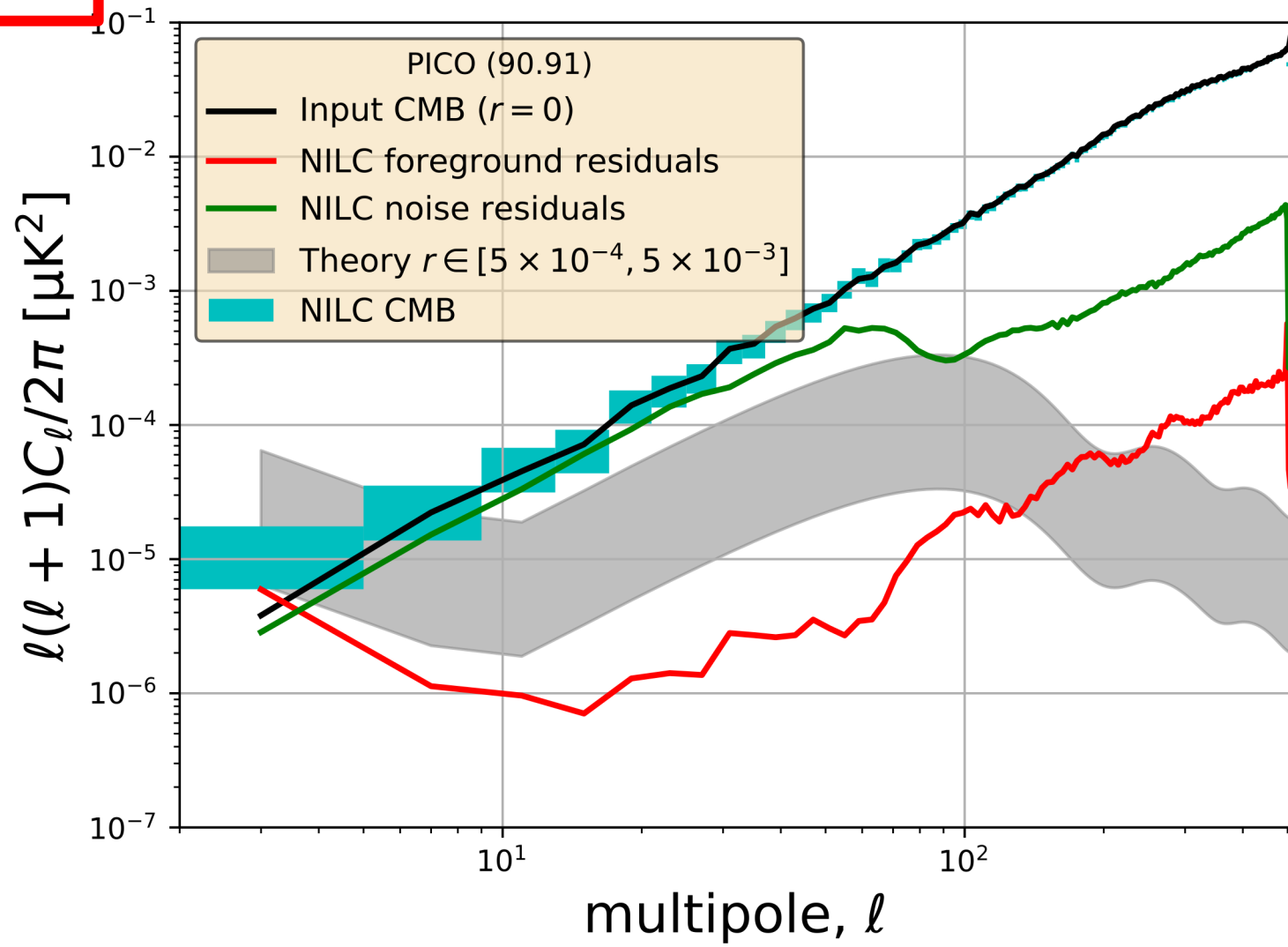
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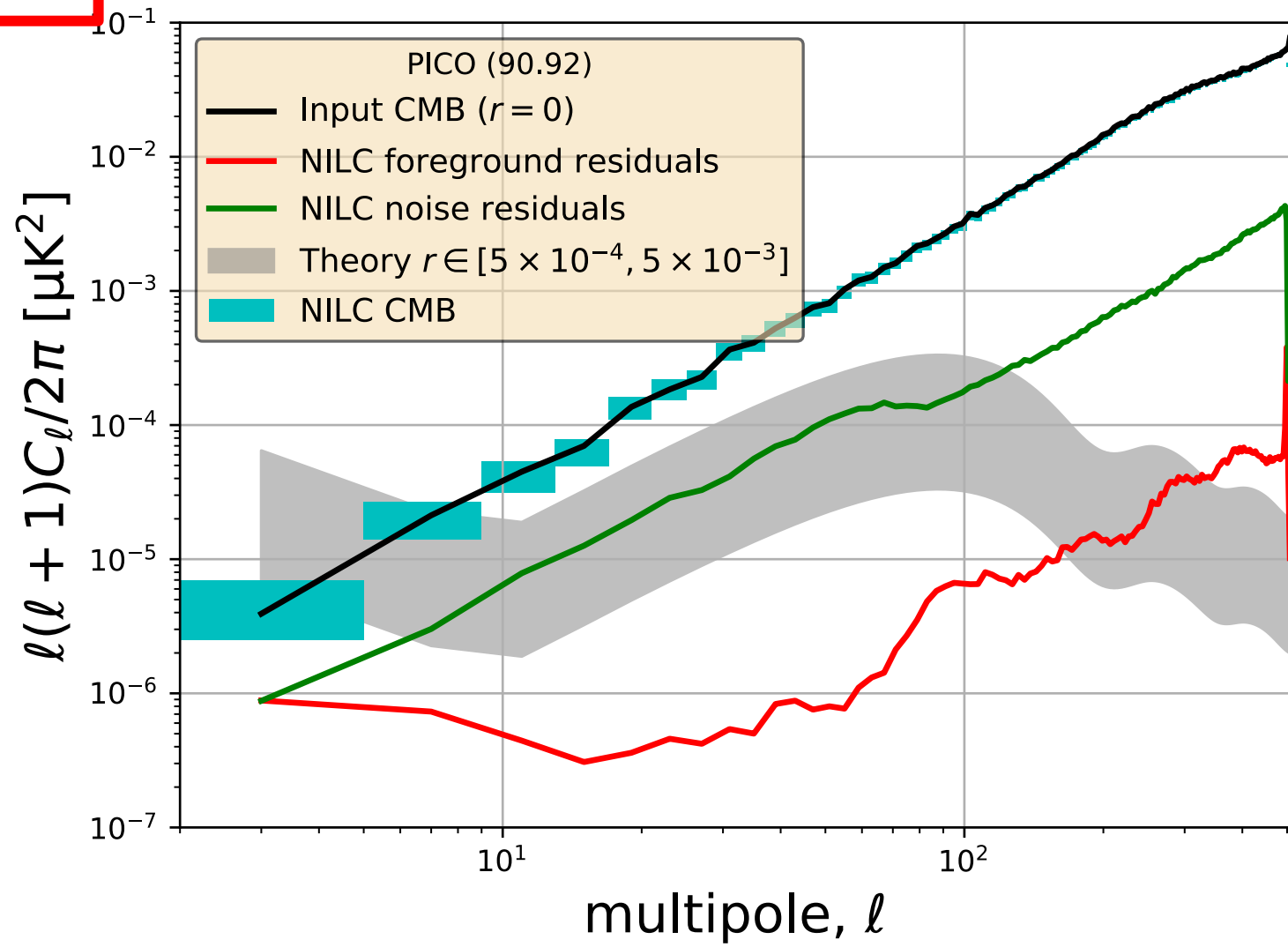


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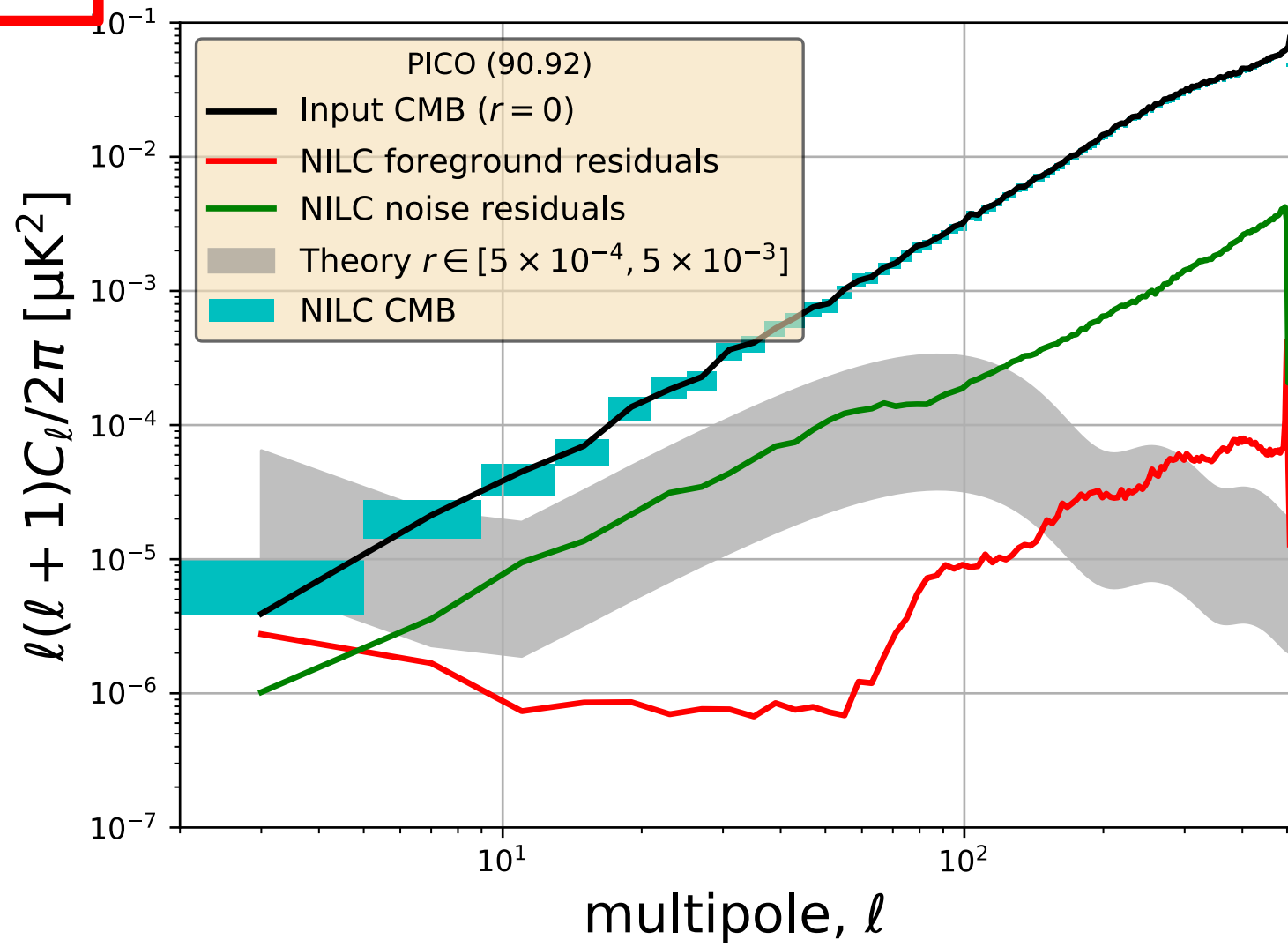
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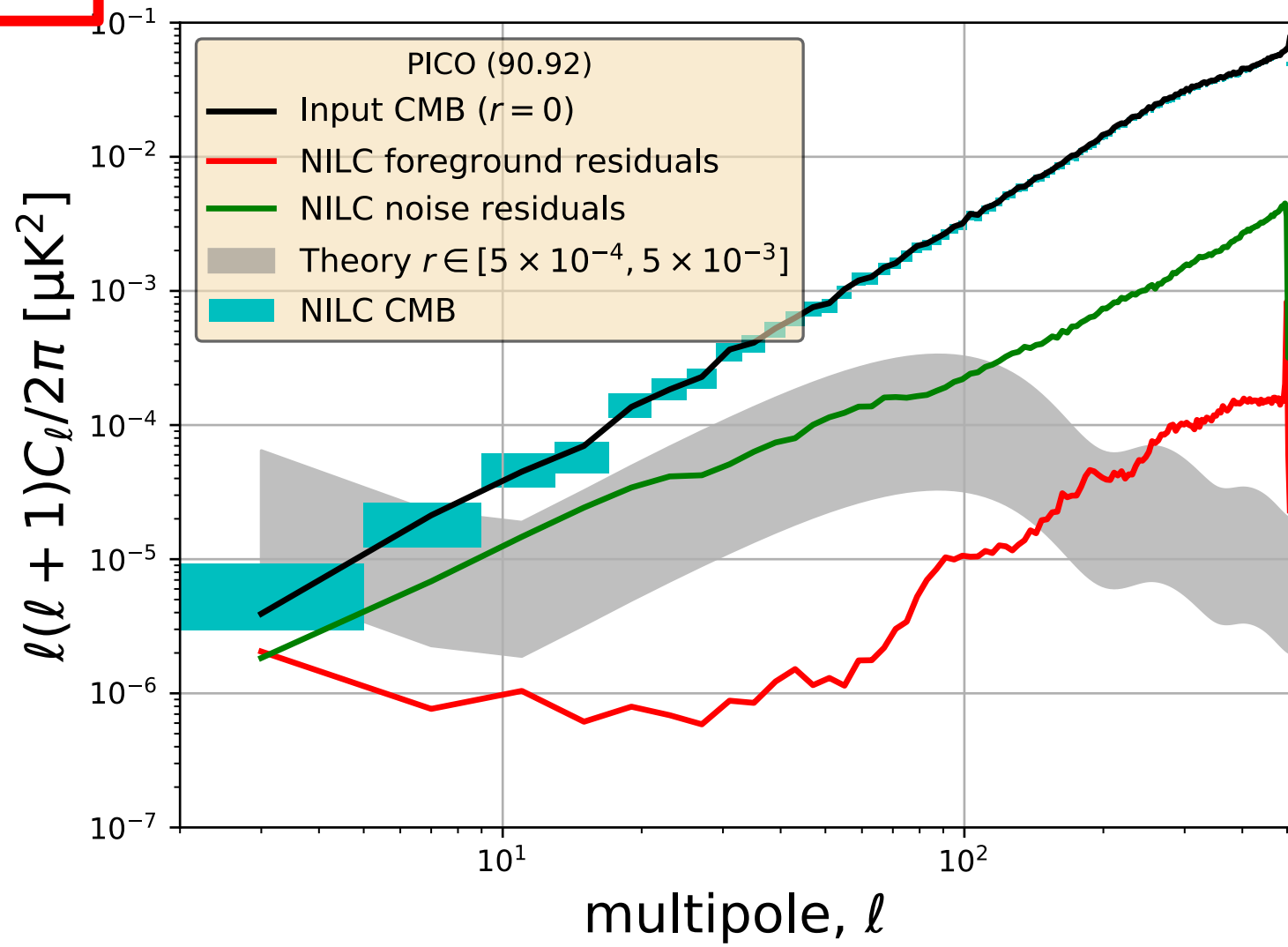
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MASTER

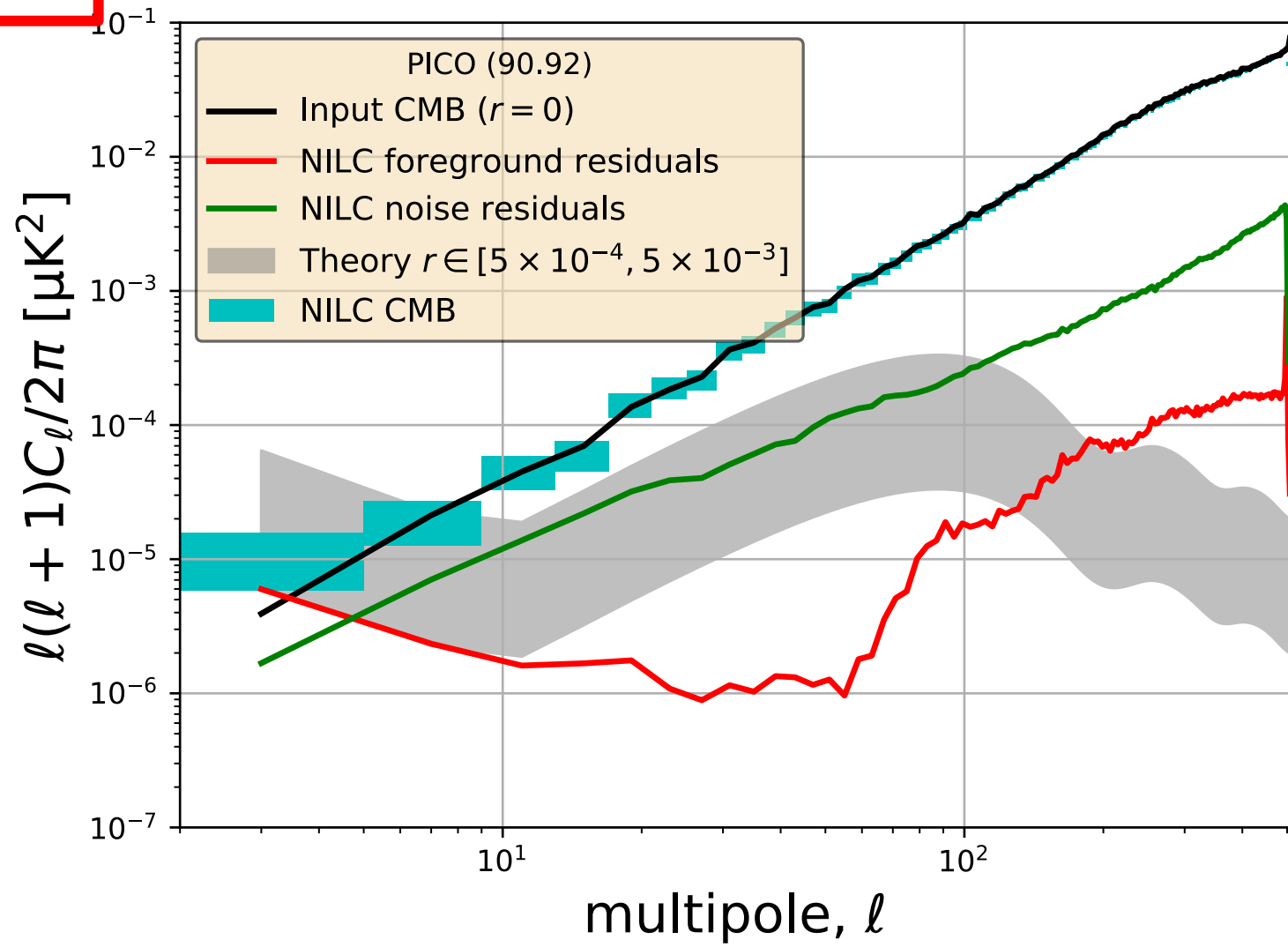
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Descope  
43-462 GHz

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10 realizations

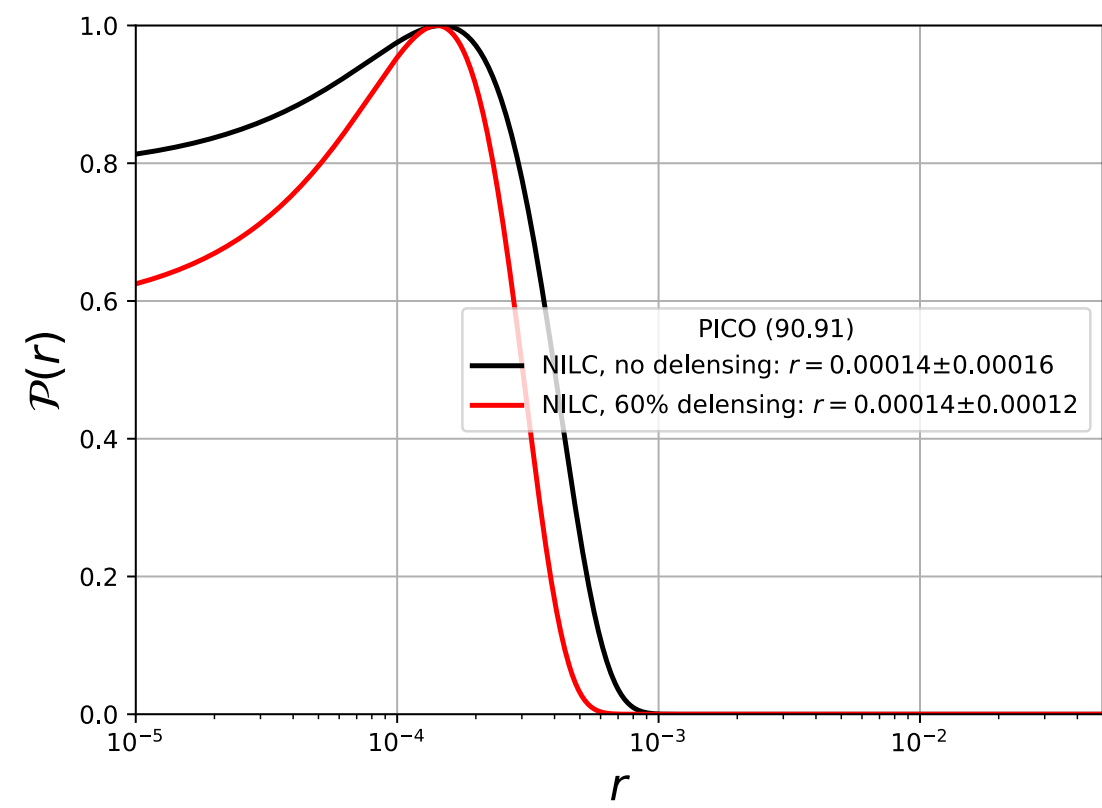


MASTER  
 $f_{\text{sky}} = 50\%$   
Binning:  $\Delta\ell = 4$

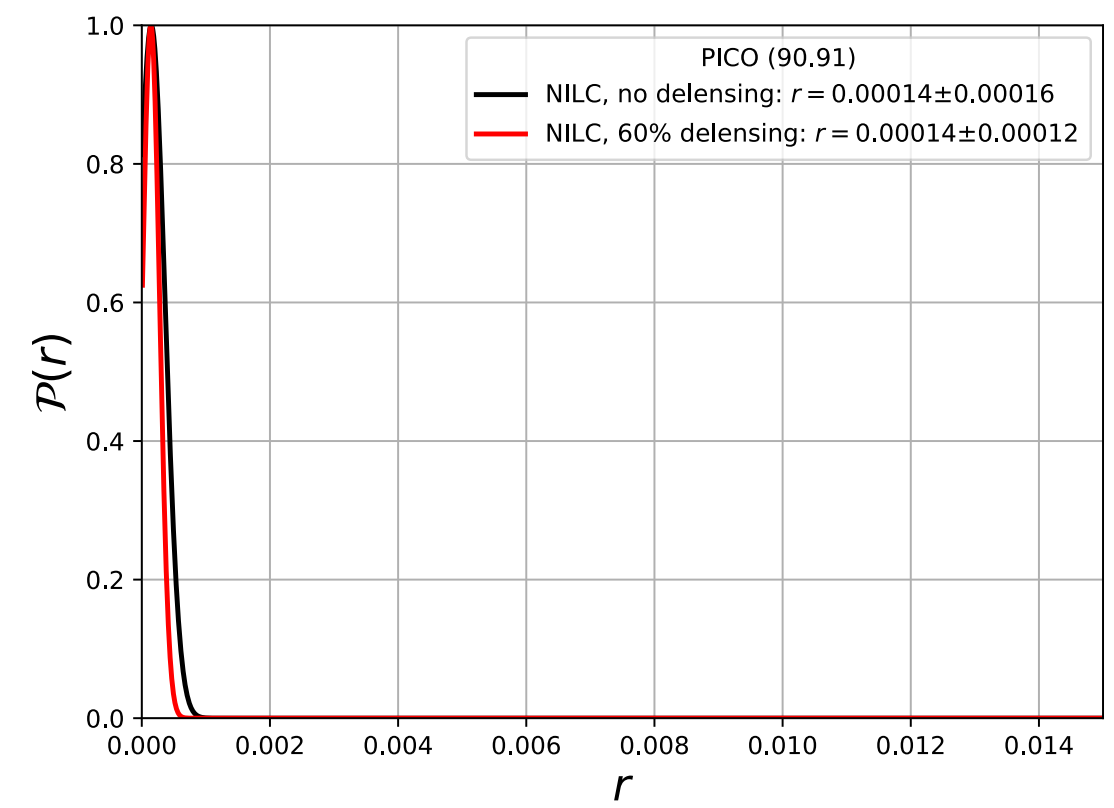
Baseline  
21-800 GHz

90.91,  $r = 0$   
NILC

Logarithmic scale



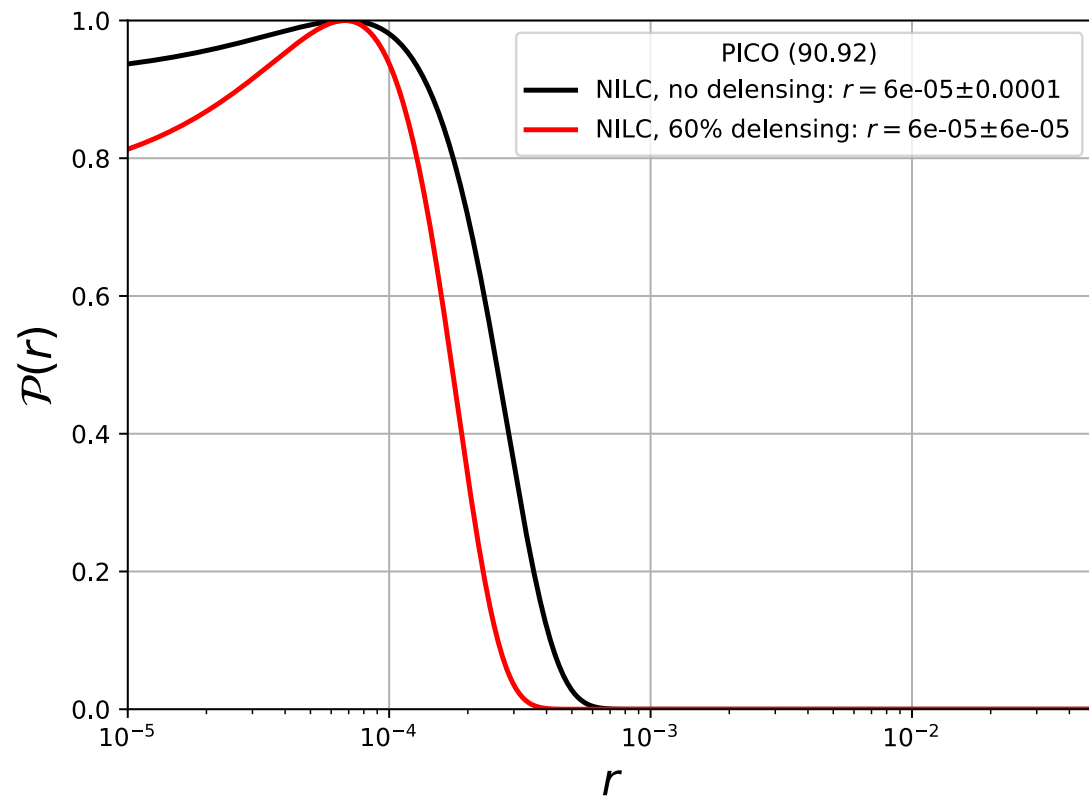
Linear scale



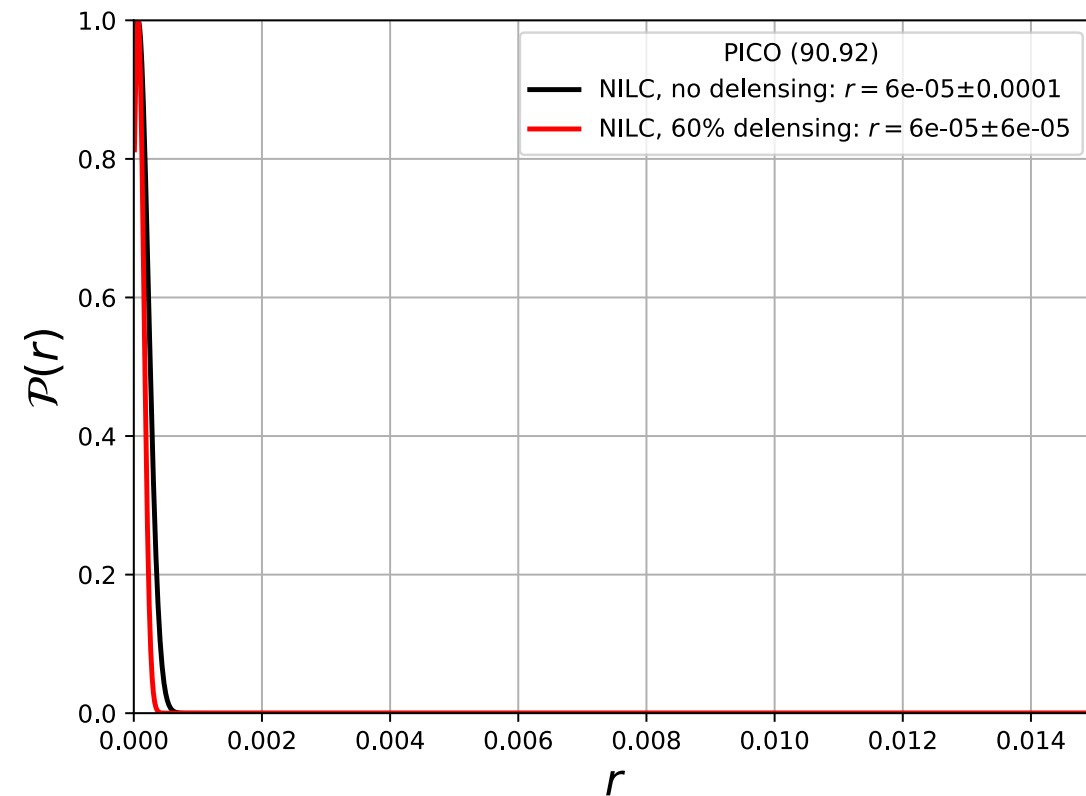
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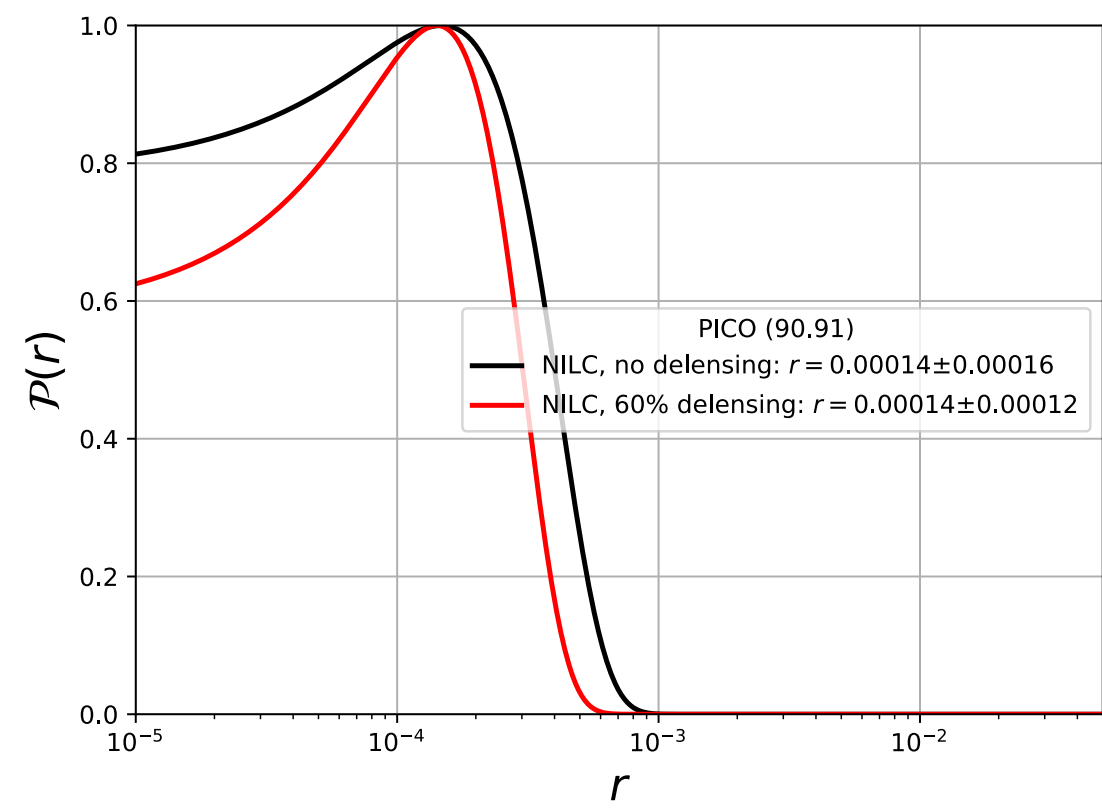
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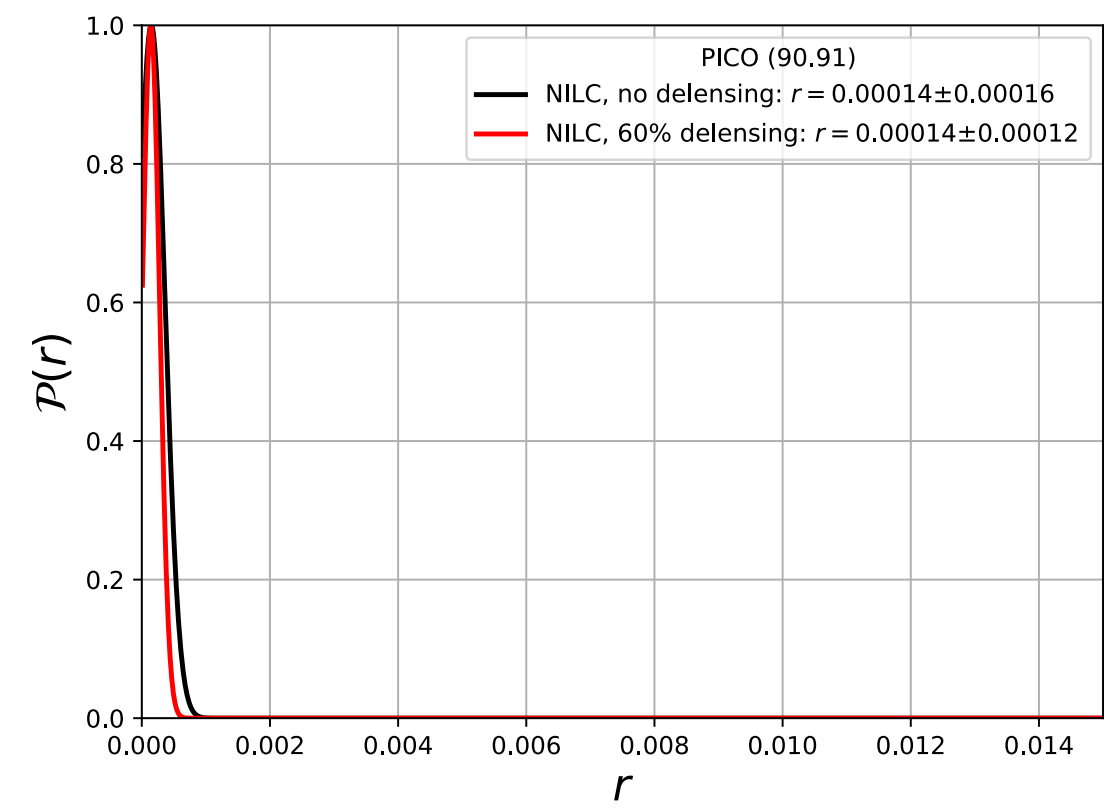
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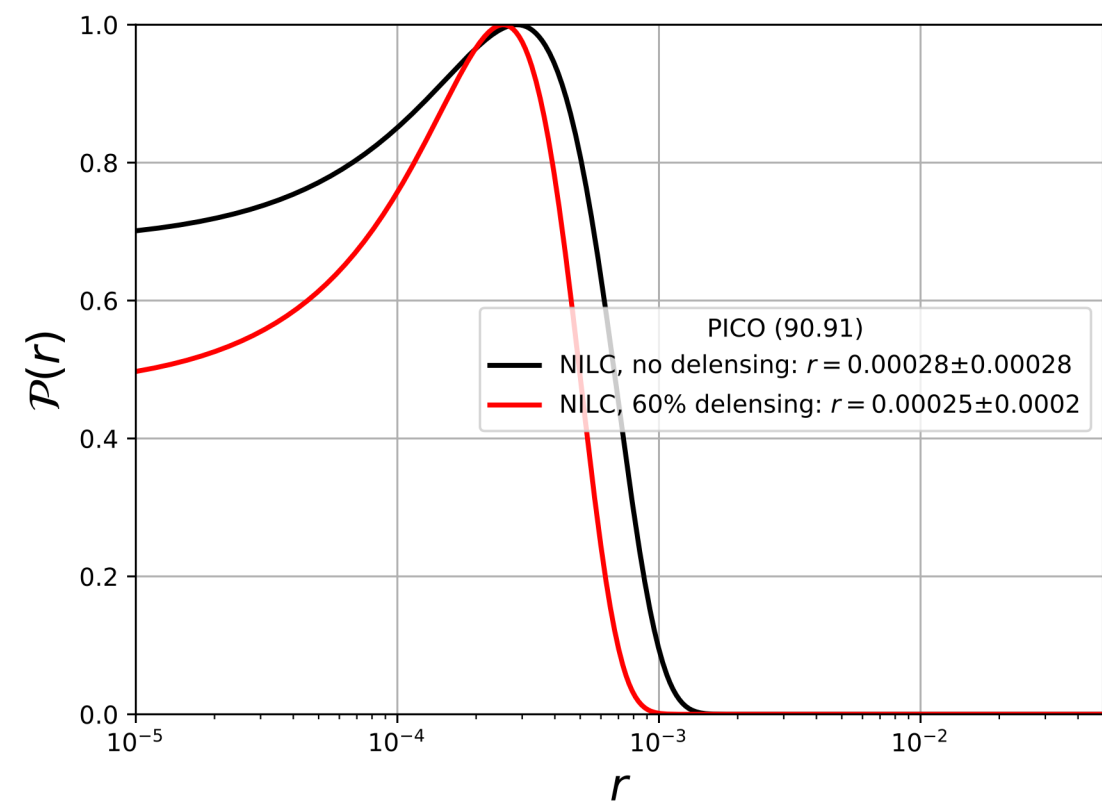
Linear scale



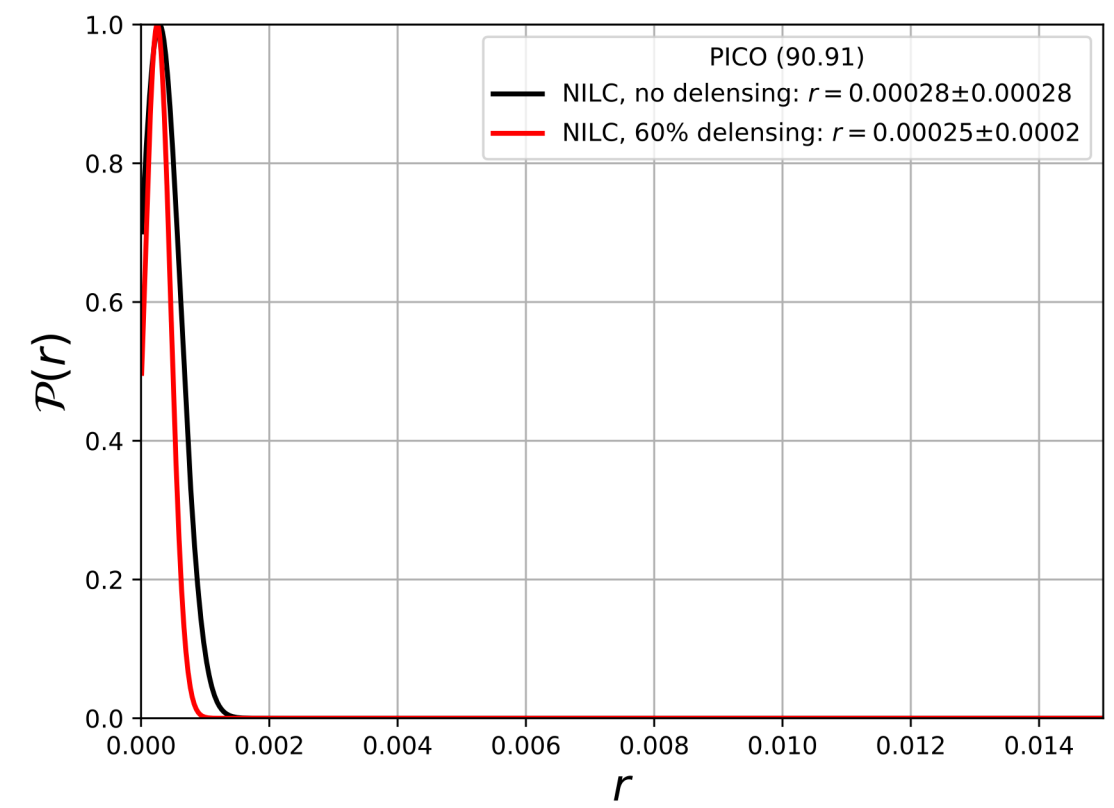
w/o LF  
43-800 GHz

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Logarithmic scale



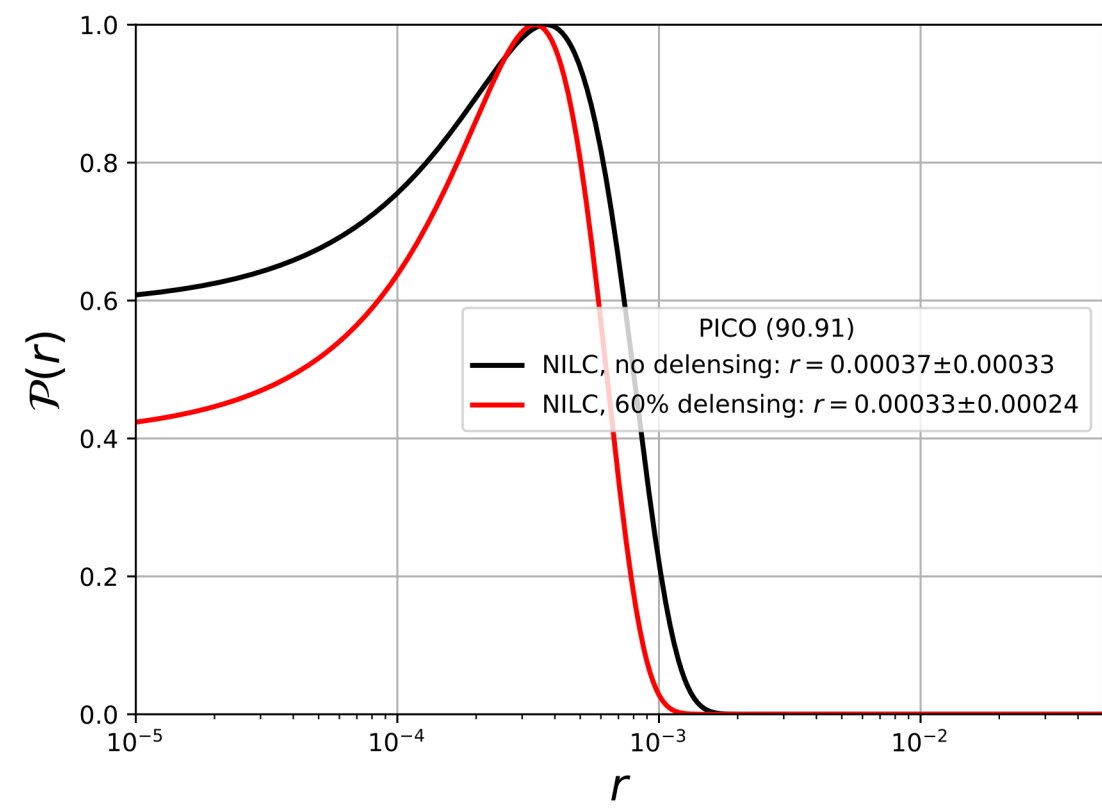
Linear scale



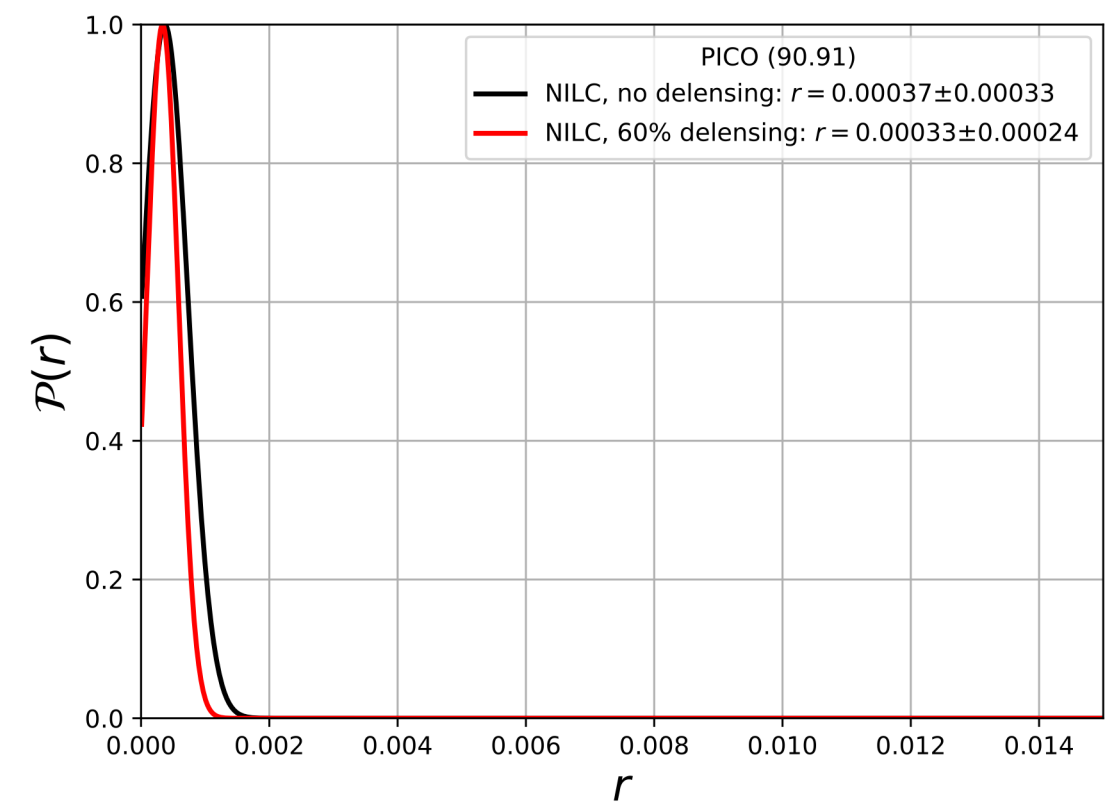
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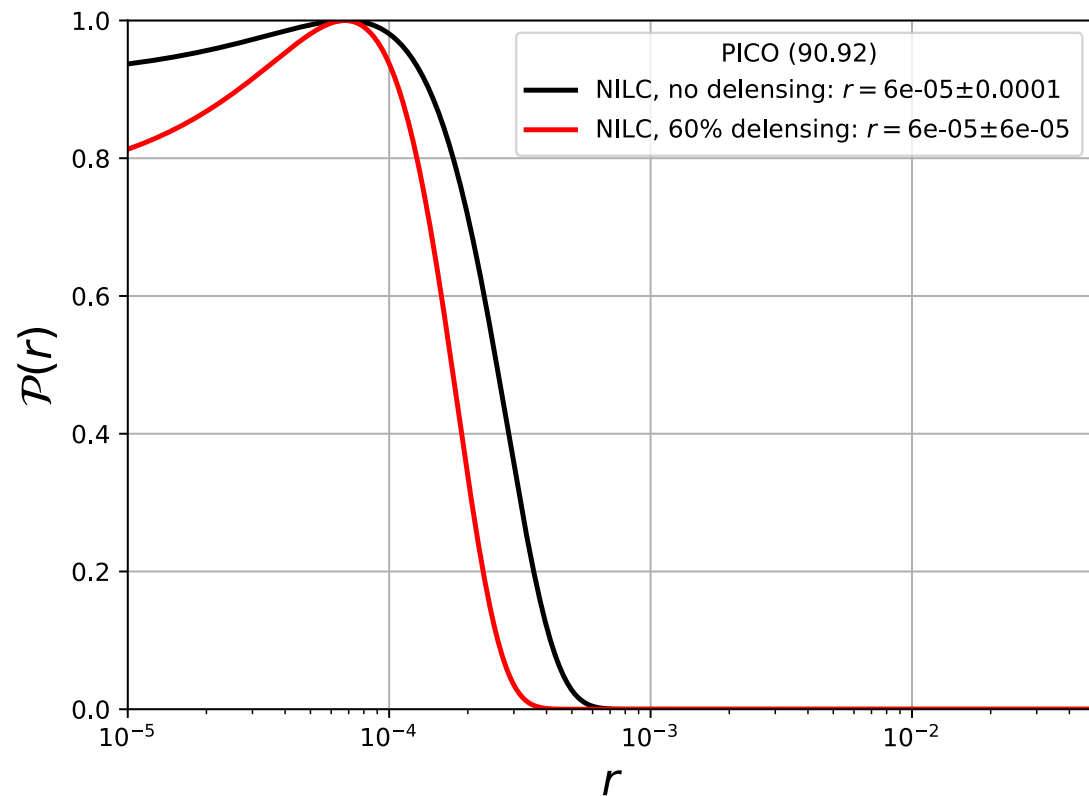
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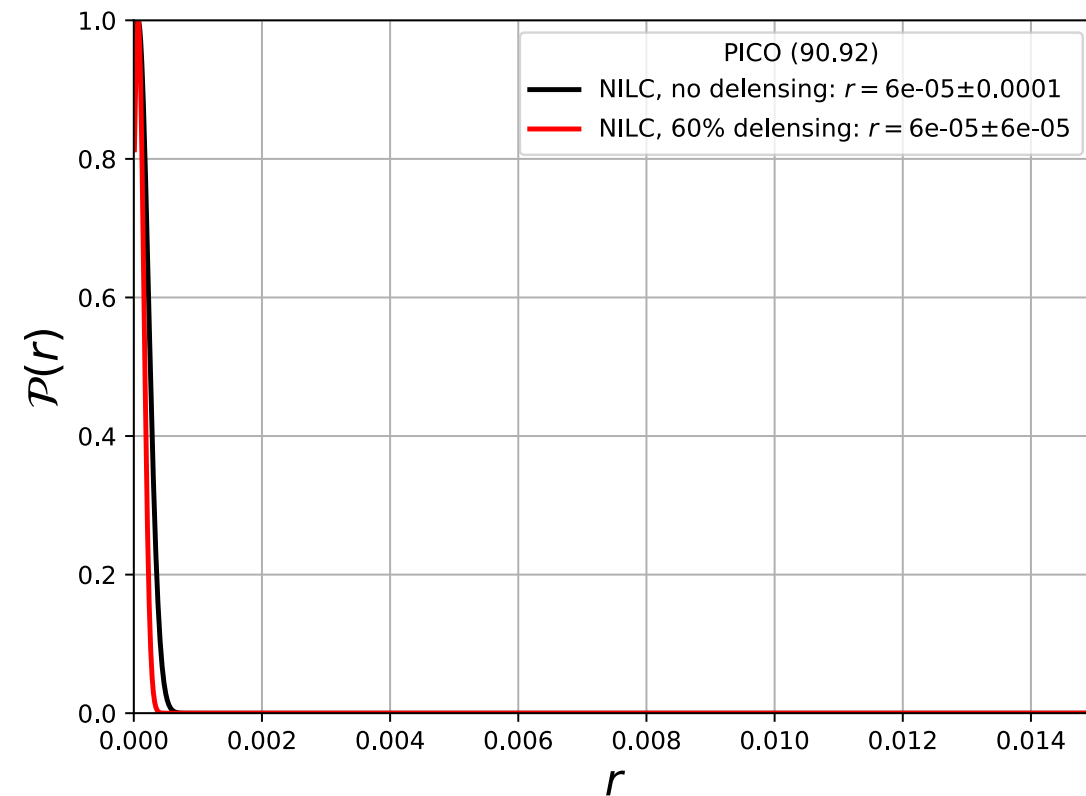
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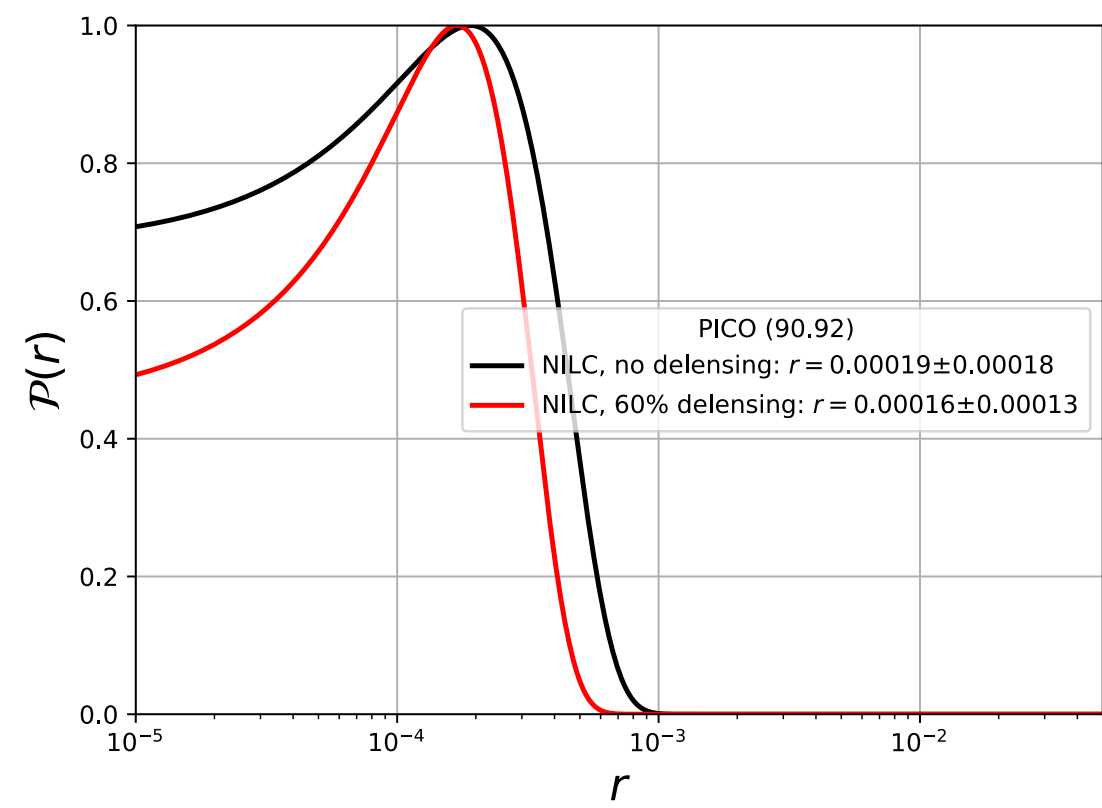
Linear scale



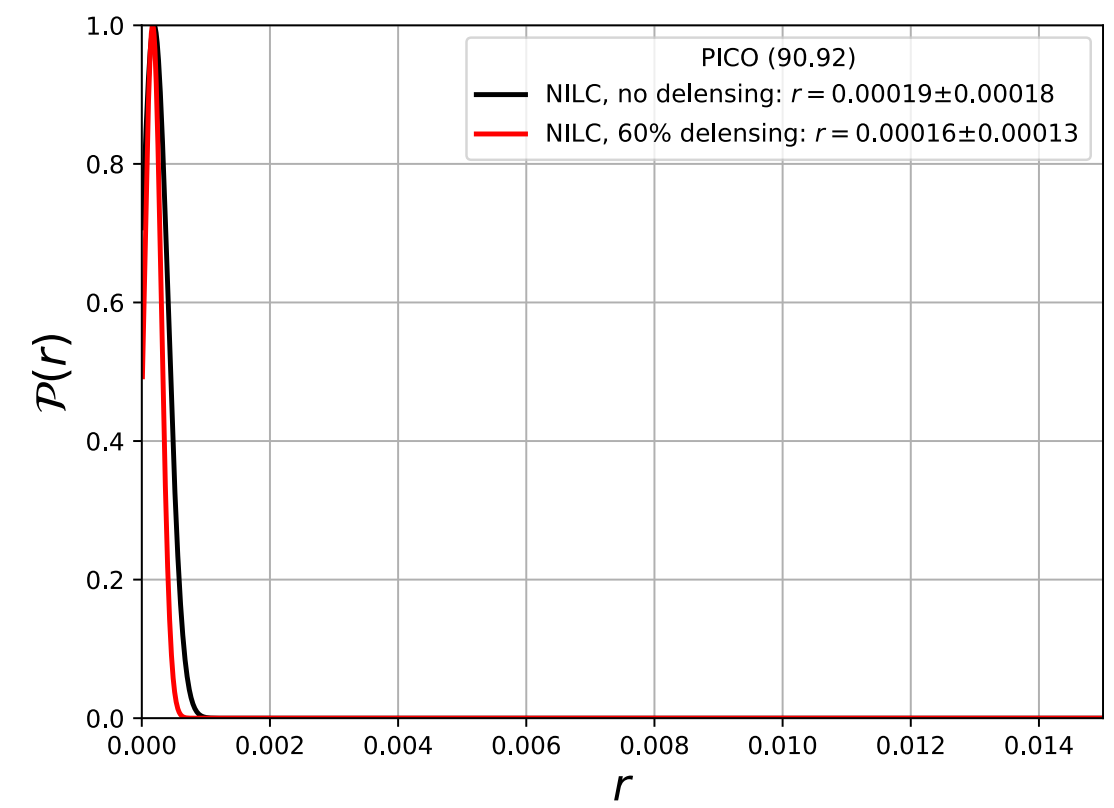
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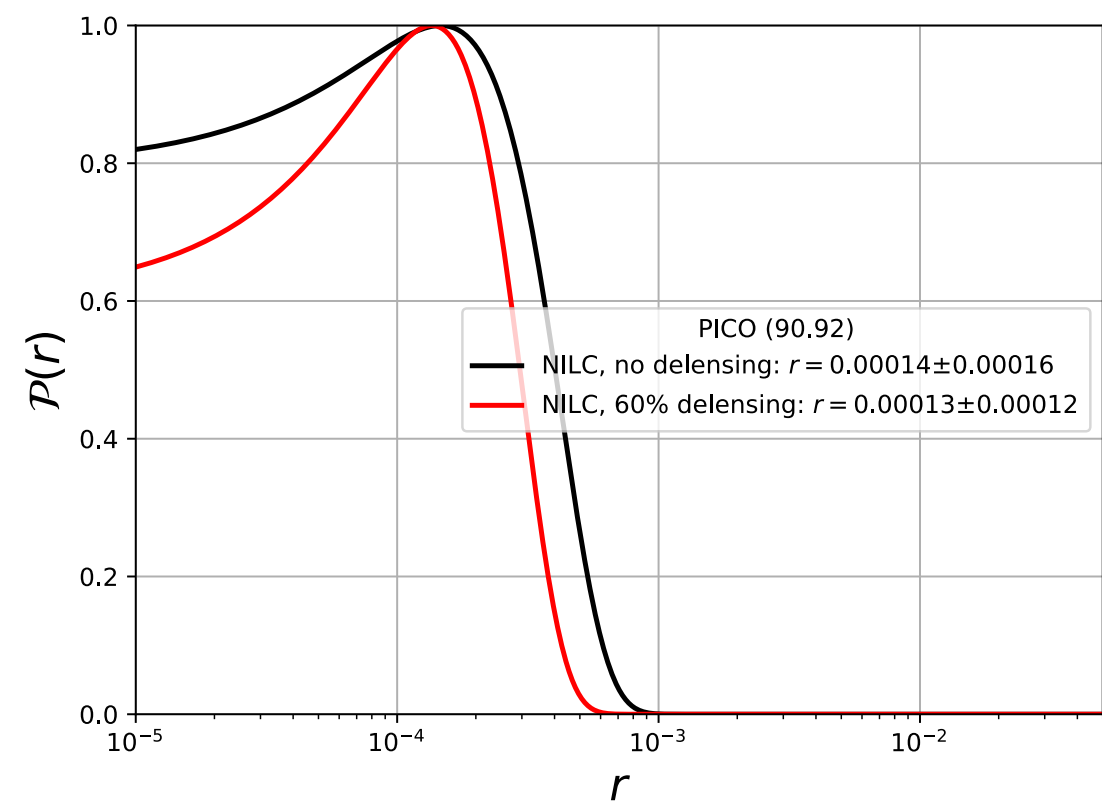
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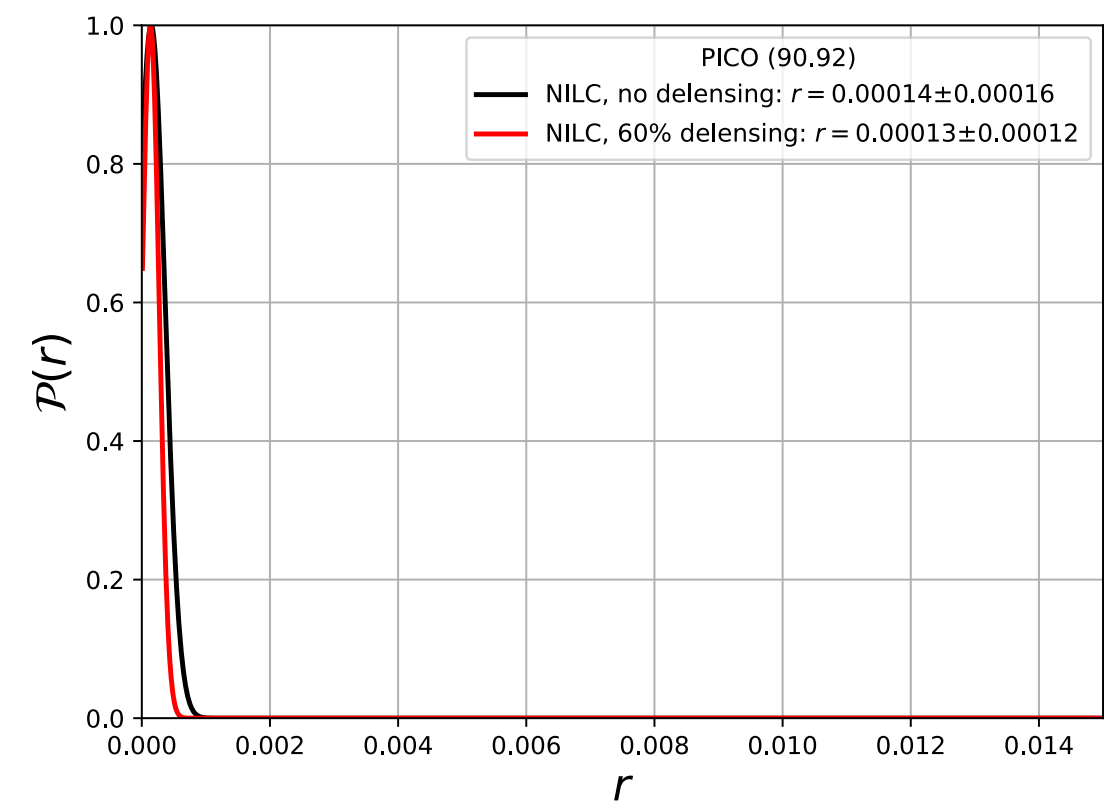
w/o HF  
21-462 GHz

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NILC

Logarithmic scale



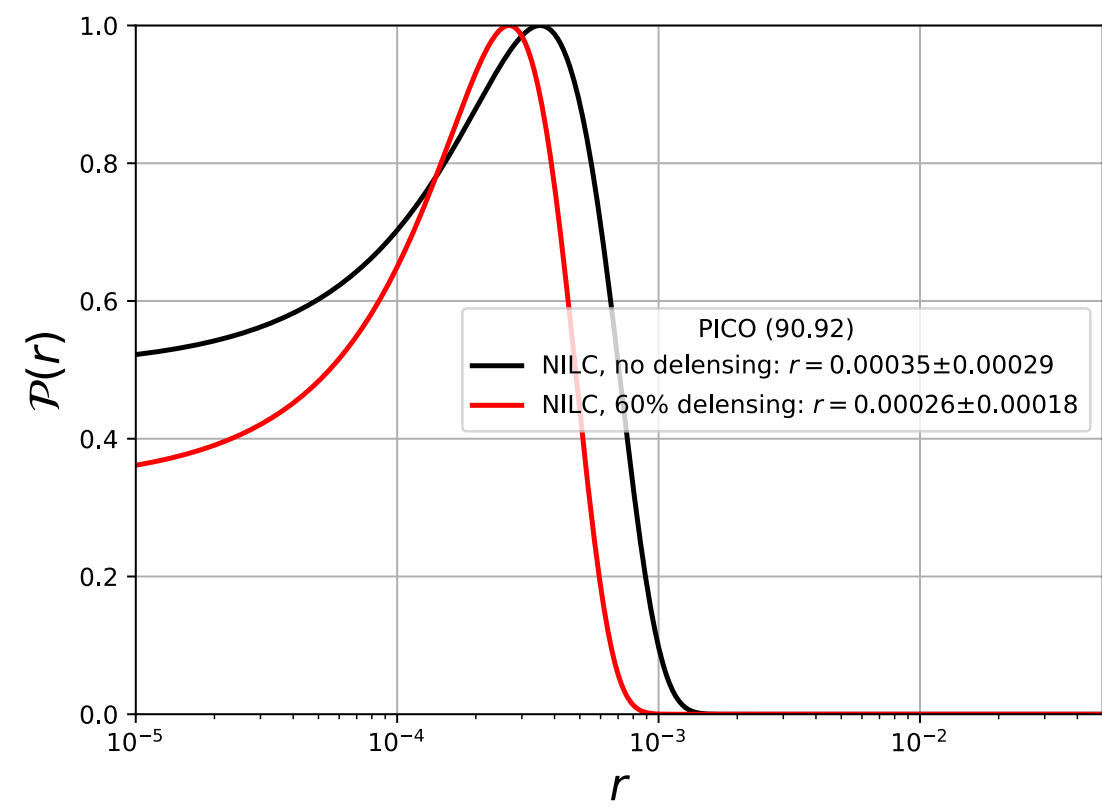
Linear scale



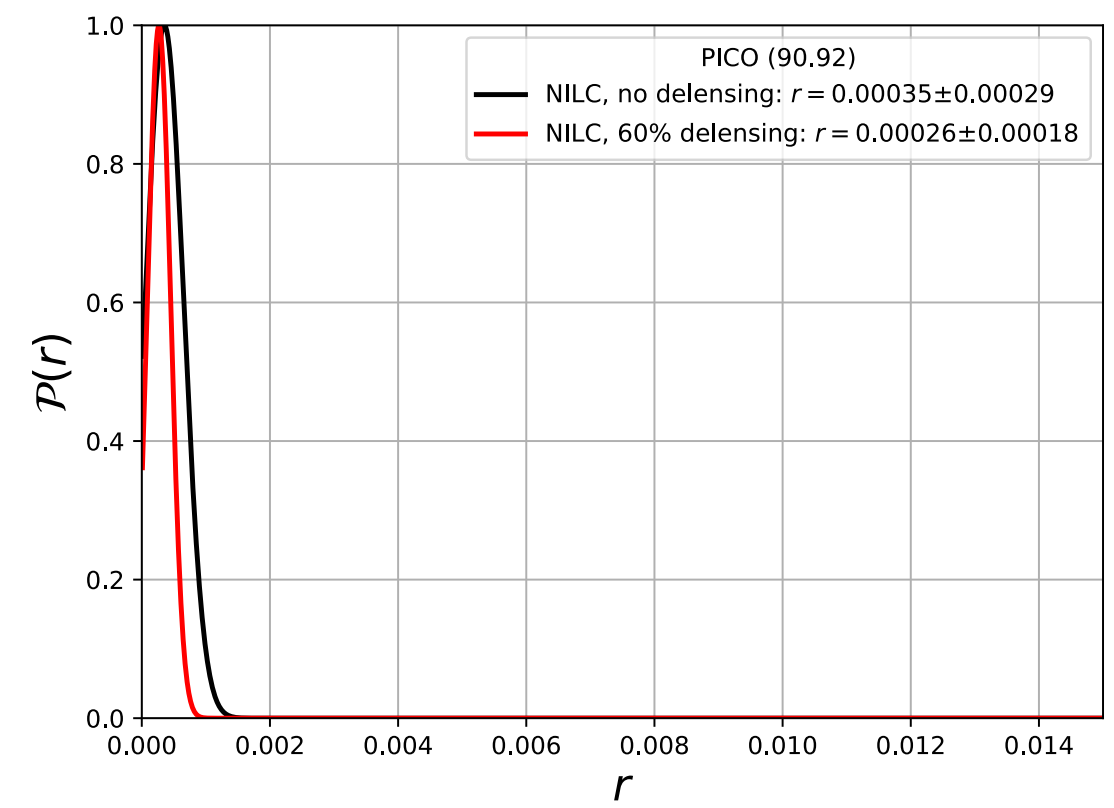
Descope  
43-462 GHz

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NILC

Logarithmic scale



Linear scale

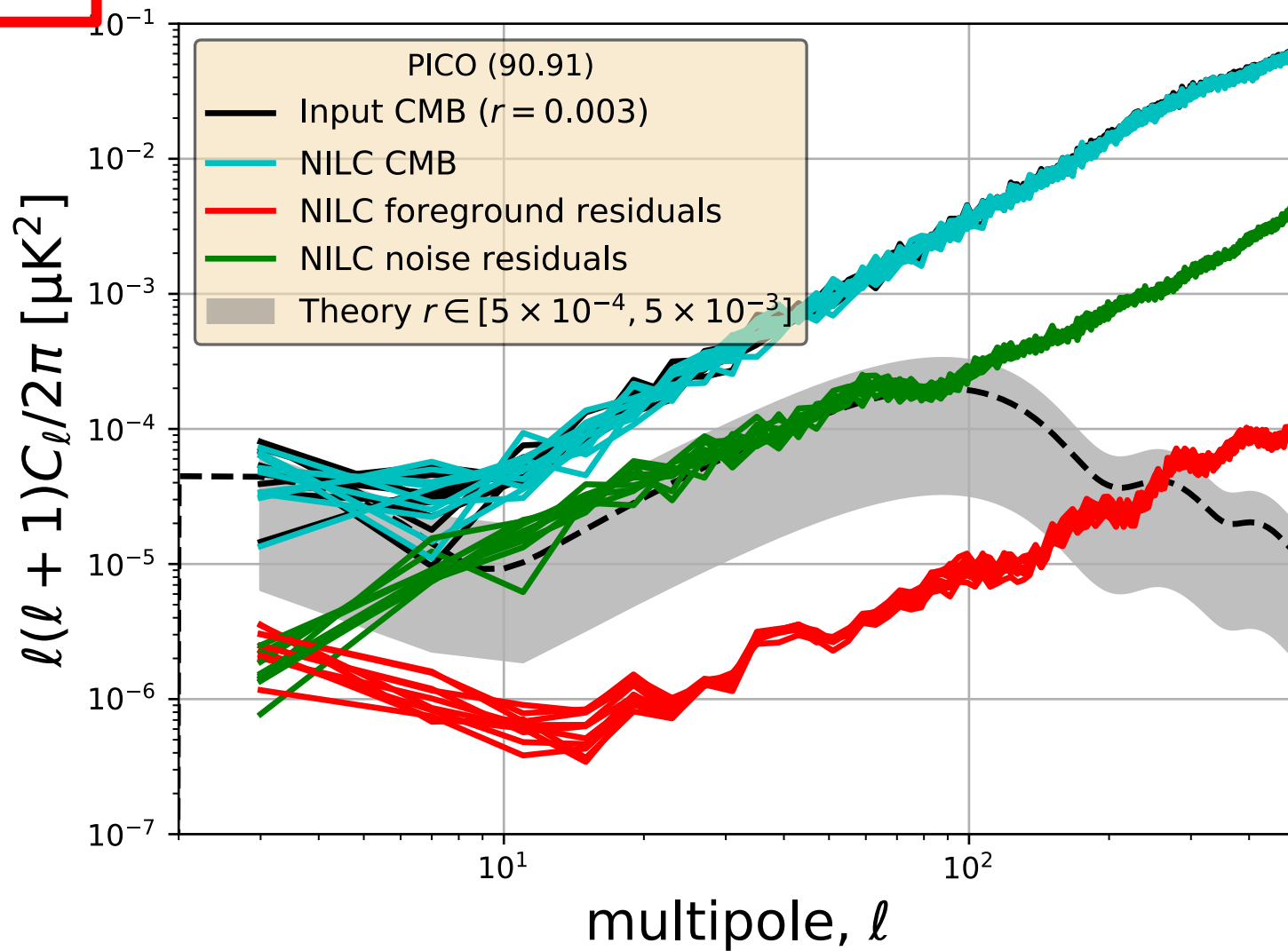


$$r = 0.003$$

90.91 & 90.92

Baseline  
21-800 GHz

90.91,  $r = 0.003$   
NILC



10 realizations

MASTER

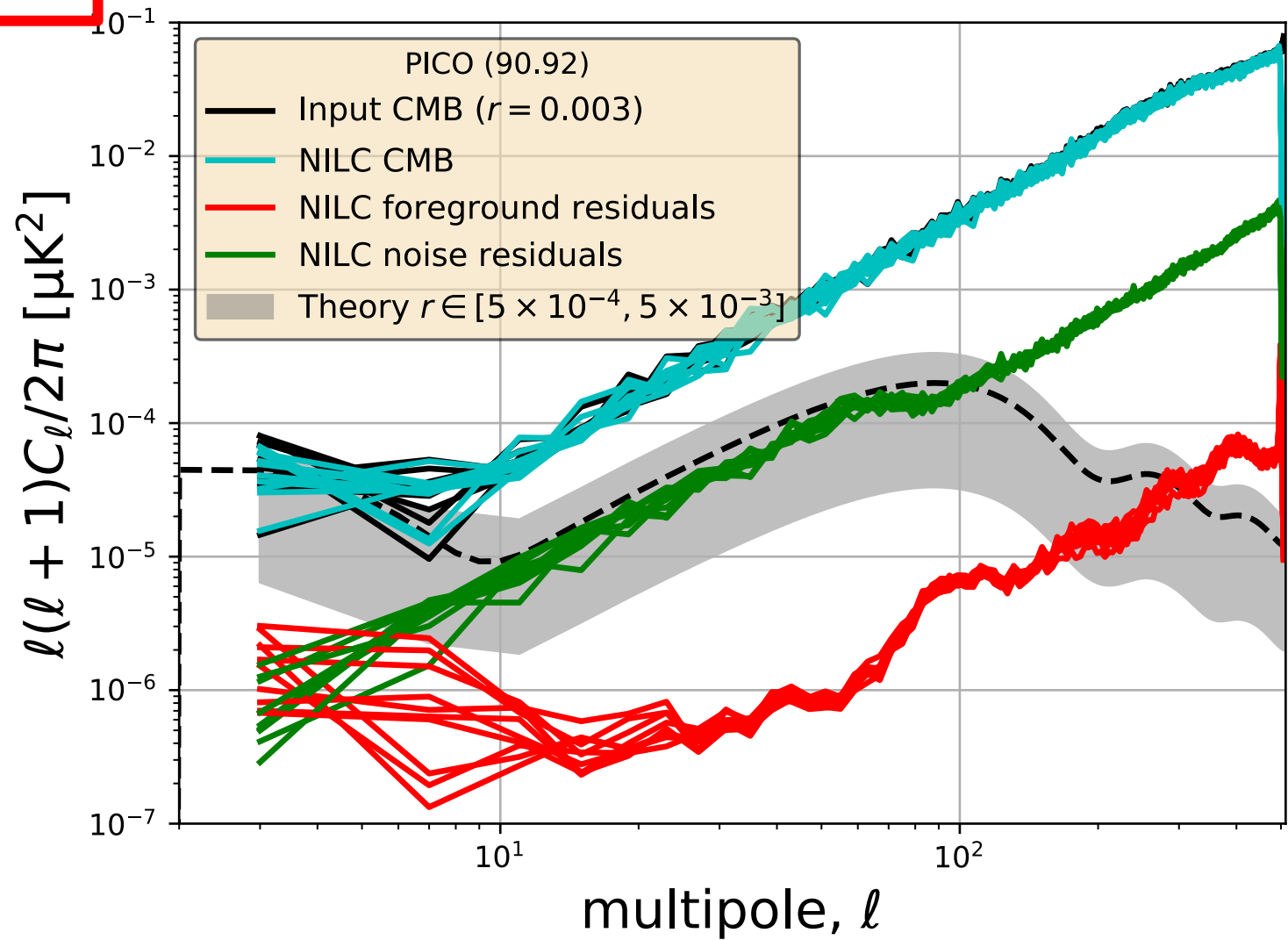
$f_{\text{sky}} = 50\%$

Binning:  $\Delta\ell = 4$

Baseline  
21-800 GHz

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NILC

10 realizations

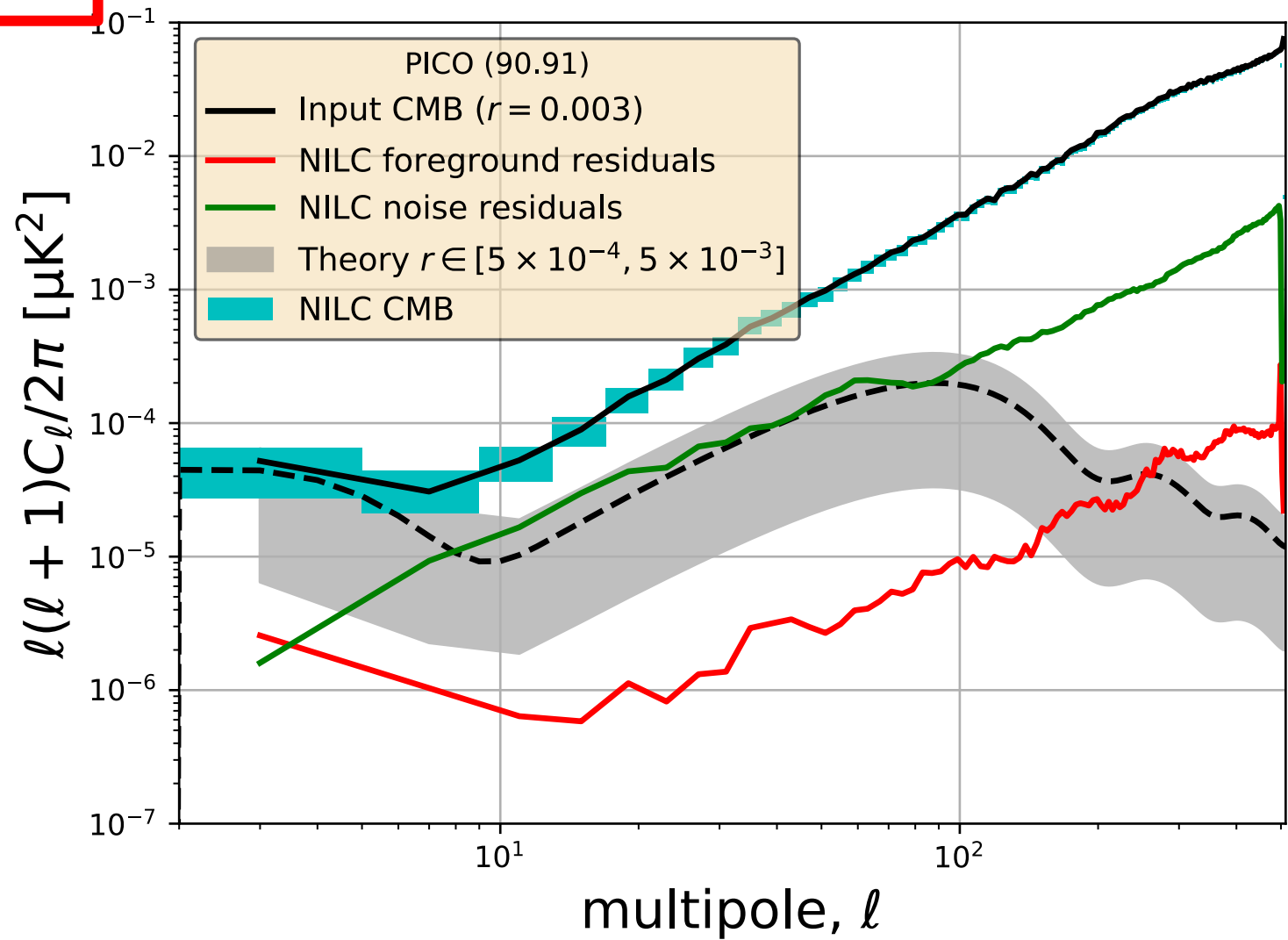


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NILC

10 realizations

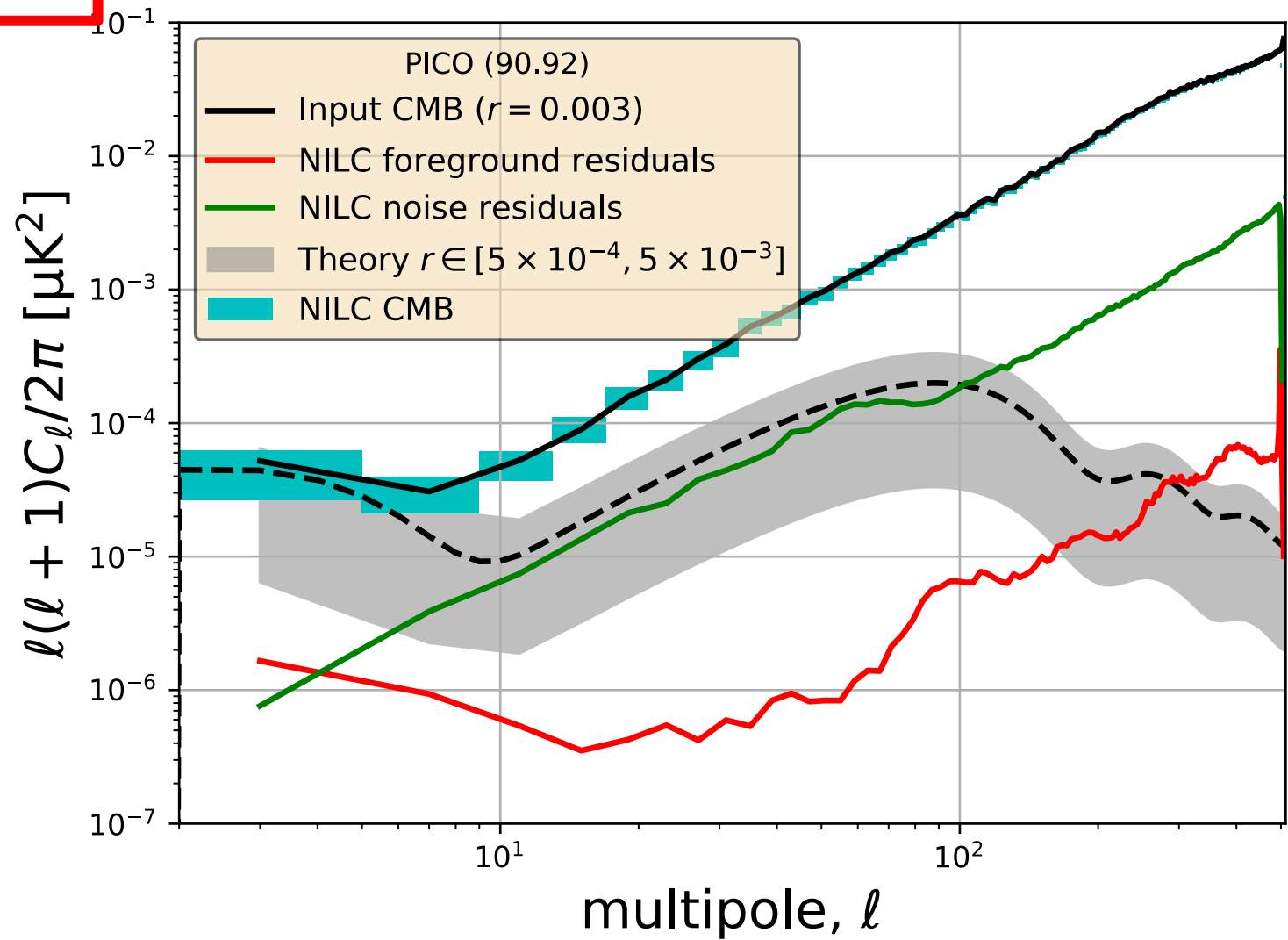


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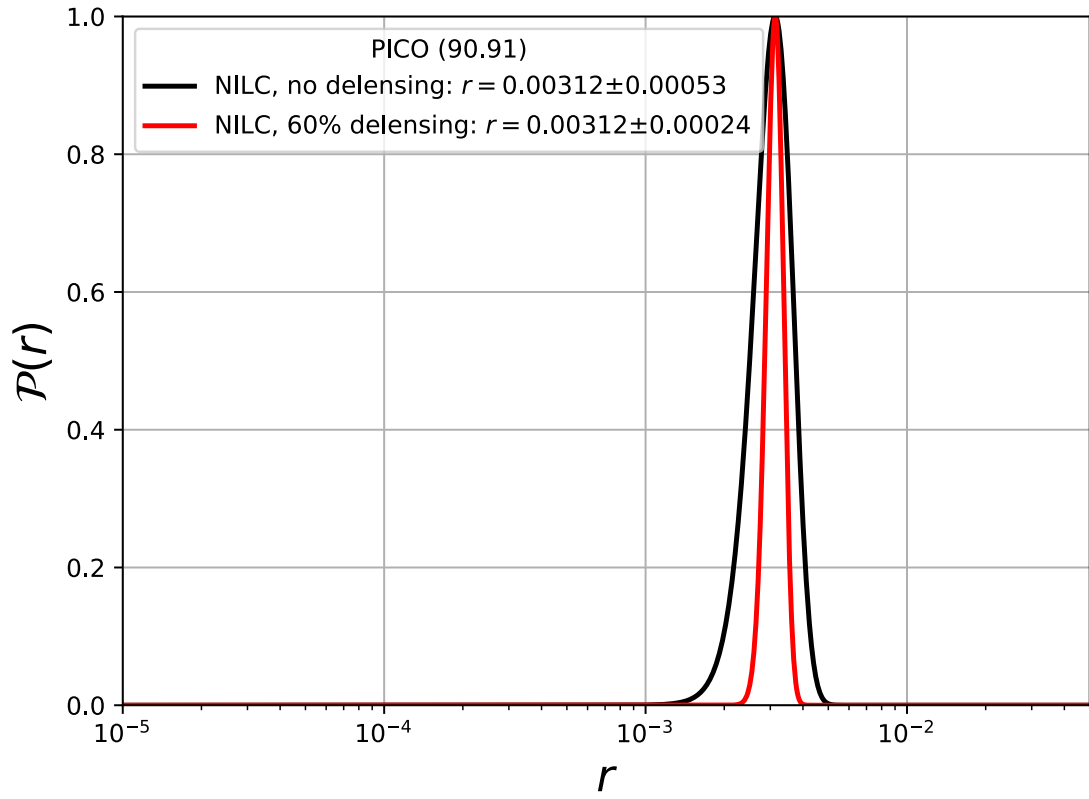


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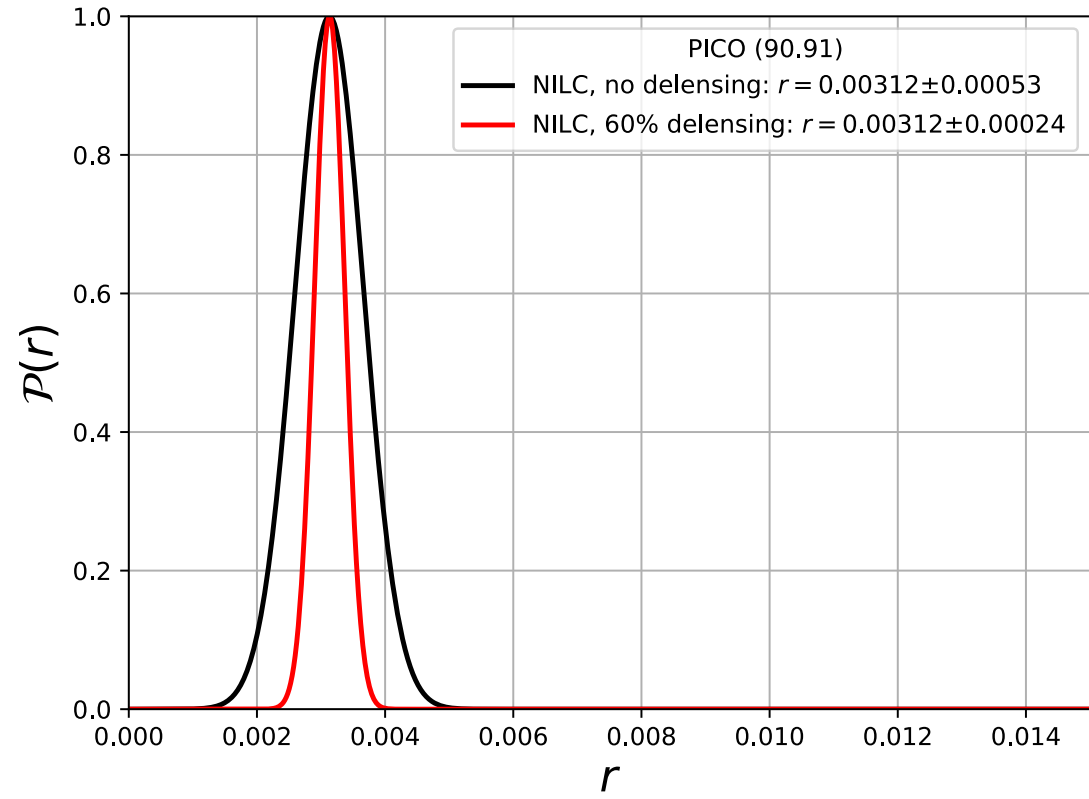
Baseline  
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NILC

*Logarithmic scale*



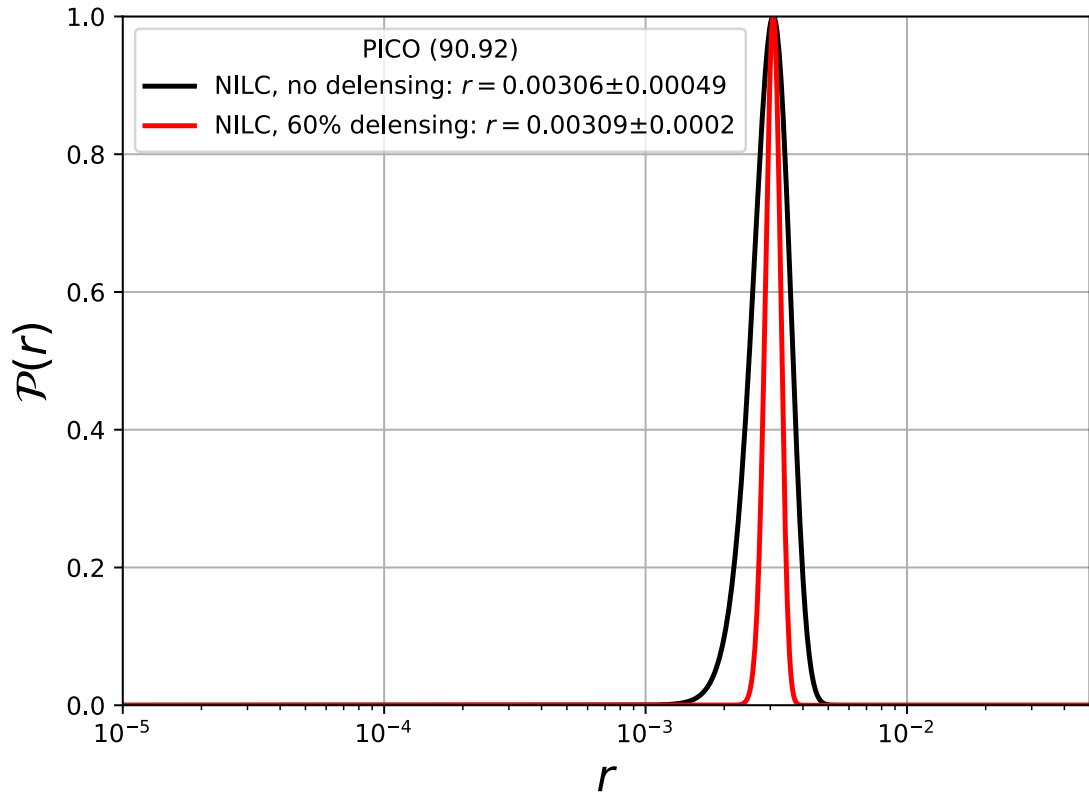
*Linear scale*



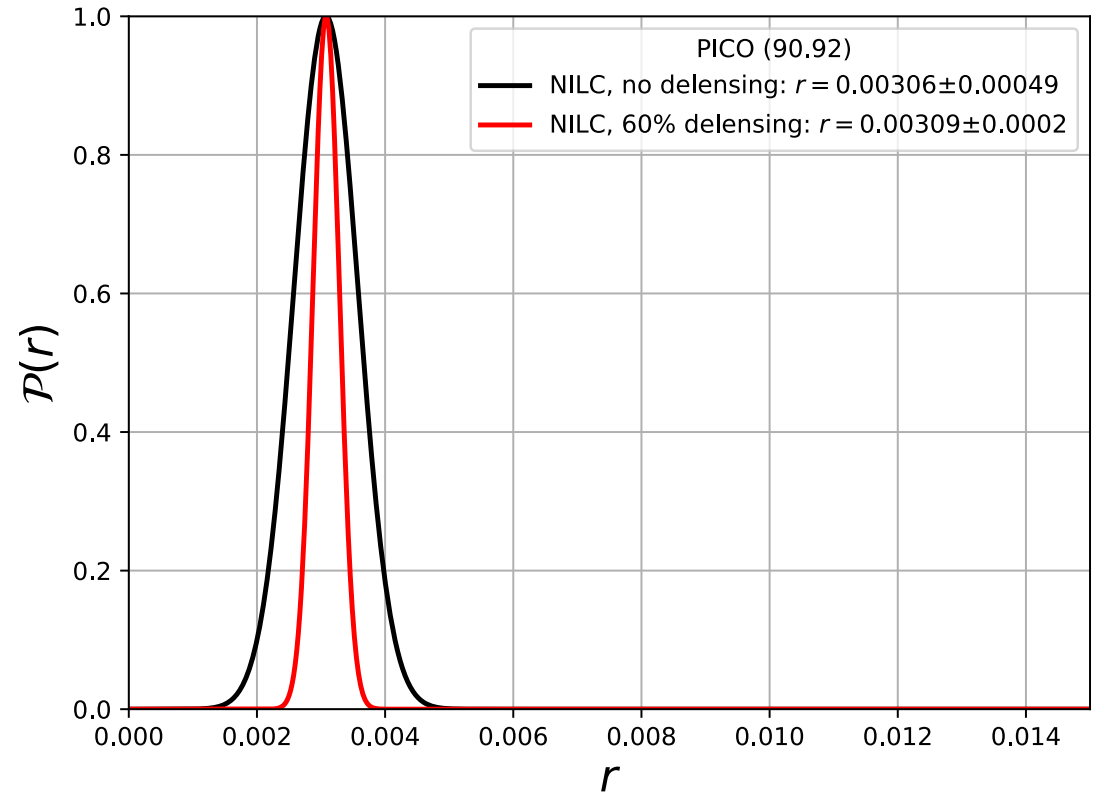
Baseline  
21-800 GHz

90.92,  $r = 0.0003$   
NILC

*Logarithmic scale*



*Linear scale*



# Note on the likelihood

The component separation exercise has been performed on sky maps with full lensing contamination.

Suppose that PICO can perform e.g. 60% delensing, then the fraction of residual lensing power will be  $A_L = 0.4$  after delensing.

Now for the  $r$  forecasts, we do the following shortcut to account for “delensing”:

- ❑  $C_\ell^{BB,NILC}$  is corrected for the residual noise bias and the residual lensing bias:

$$C_\ell^{\text{CMB}} + C_\ell^{\text{fgds}} = C_\ell^{BB,NILC} - C_\ell^{\text{noise}} - A_L C_\ell^{\text{lens}}$$

- ❑ Build a simple Gaussian likelihood to fit  $r$  only:

$$-2 \ln \mathcal{L}(r) = \sum_{\ell=2}^{\ell_{\max}} \left( C_\ell^{\text{CMB}} + C_\ell^{\text{fgds}} - r C_\ell^{\text{prim}}(r=1) \right) M_{\ell\ell'}^{-1} \left( C_\ell^{\text{CMB}} + C_\ell^{\text{fgds}} - r C_\ell^{\text{prim}}(r=1) \right)$$

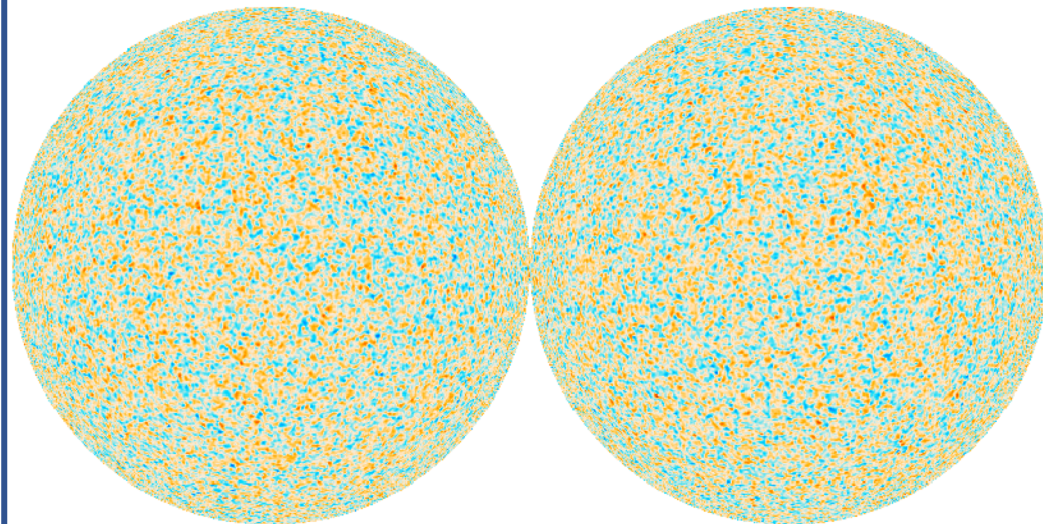
- ❑ The covariance matrix includes cosmic/sample variance of residual lensing signal, residual foregrounds and residual noise (and cross-terms):

$$M_{\ell\ell} = \frac{2}{(2\ell+1)f_{\text{sky}}} \left( C_\ell^{BB,NILC} - (1 - A_L) C_\ell^{\text{lens}} \right)^2 = \frac{2}{(2\ell+1)f_{\text{sky}}} \left( C_\ell^{\text{CMB}} + A_L C_\ell^{\text{lens}} + C_\ell^{\text{fgds}} + C_\ell^{\text{noise}} \right)^2$$

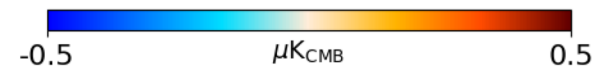
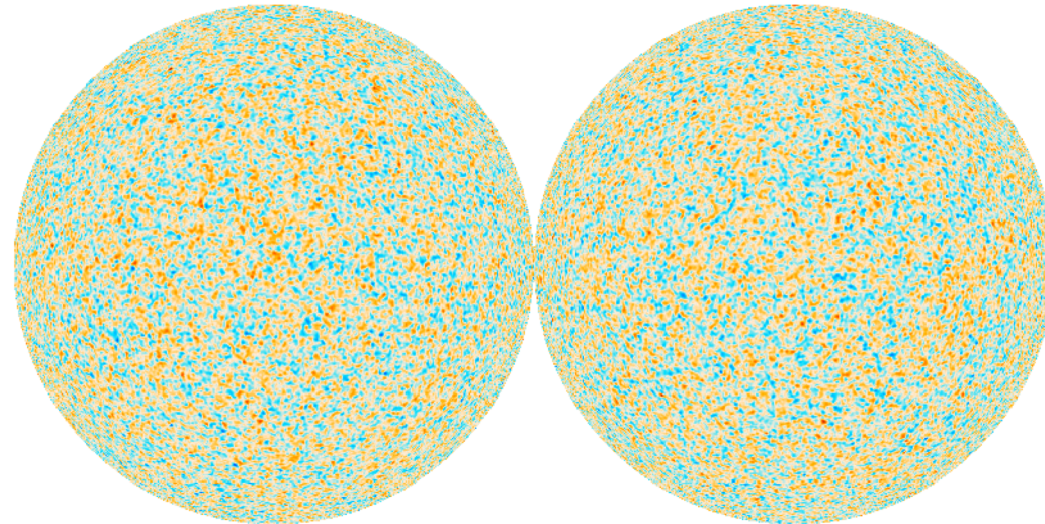
 Expected residual lensing cosmic variance

*Backup*

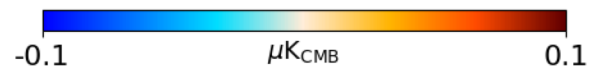
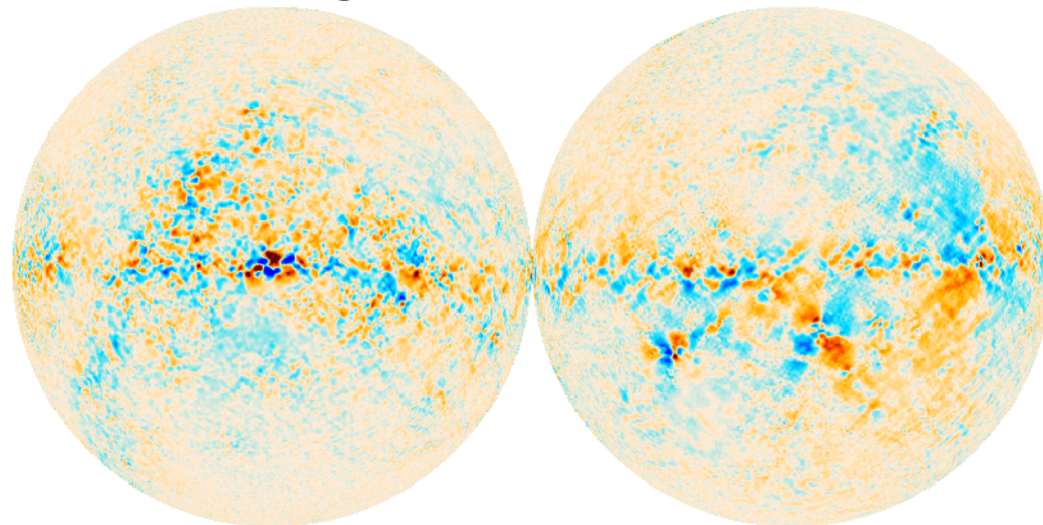
INPUT CMB B-mode  $r = 0$  (0001)



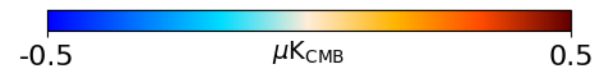
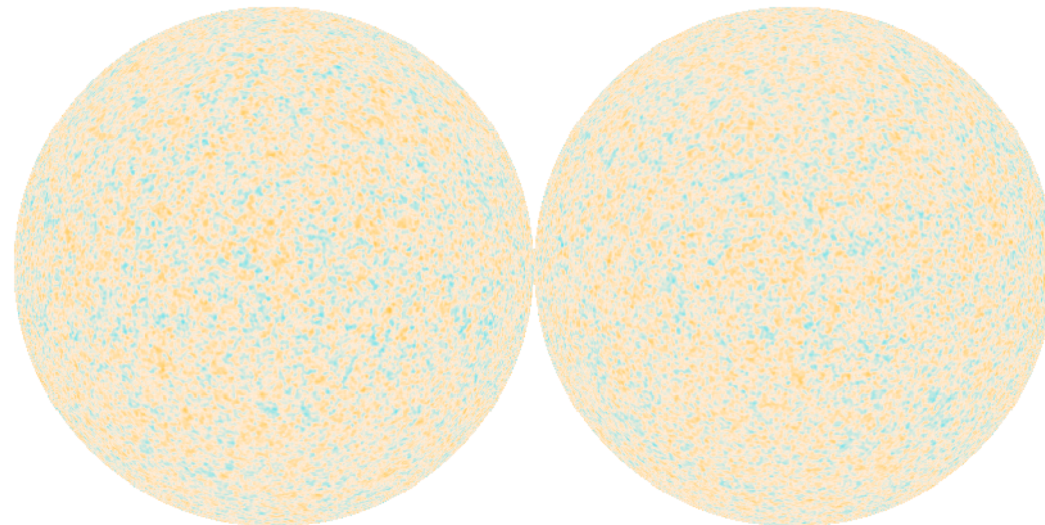
NILC CMB B-mode (0001)



NILC foreground residuals B-mode (0001)



NILC noise residuals B-mode (0001)



90.91  
 $r = 0$   
NILC