

Cosmogenic Simulations with Geant4 and FLUKA

A.N. Villano ¹ A. Empl ²

¹University of Minnesota

²University of Houston

November 11, 2011

Table of contents

- 1 The Goals and Physics
- 2 The Standard Simulations
- 3 Further Studies
- 4 Summary

what do we want?

for low background underground experiments

How many neutrons will not be able to be vetoed in an experiment?

- depends on neutron flux in cavern
- also depends on electromagnetic signatures and their time-space correlation

what are the problems?

for low background underground experiments

- measurements are difficult for a variety of reasons
 - data slow to come in
 - suitable locations hard to find
 - correlations in time and space almost impossible to measure
 - often small effects in tails of distributions can be important
-
- still there are a number of experiments with useful data (NMM,LVD,Borexino,KAMLAND...)
 - need more physics (particle transport equations or track-by-track simulation)
 - correlations difficult with particle transport so usual approach is simulation

standard software

- for cosmogenics this means studying $\mu^{+/-}$ cascades virtually, then using data to verify
- Geant4 and FLUKA are widely used for this purpose in underground science

Geant4

- <http://geant4.cern.ch>
- latest version: v4.9.4.p01
- does not yet include latest update to μ -nuclear interaction

FLUKA

- <http://www.fluka.org>
- latest version: v2011.2.6

standard software (2)

Geant4 (our version: v4.9.3+)

- open source
- particle tracking and microscopic interactions included
- modular and customizable by users

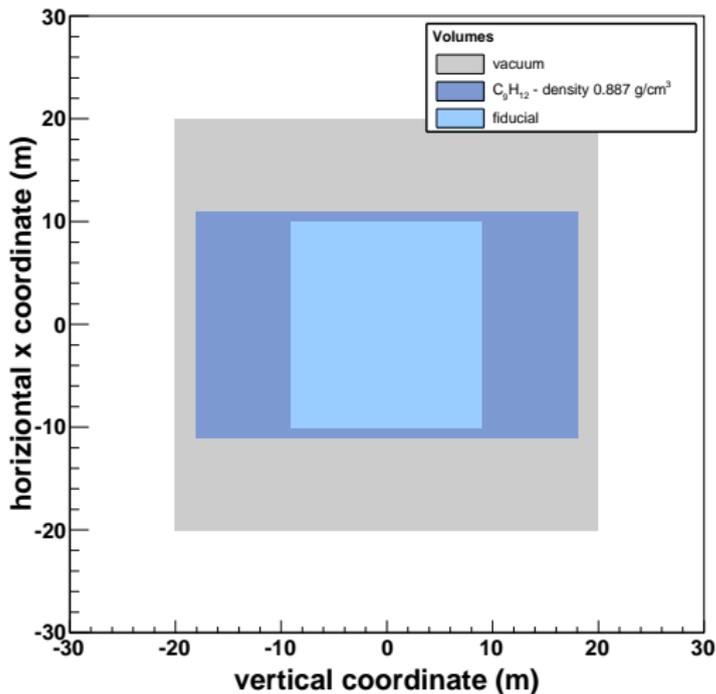
FLUKA (our version: v2011.2.5)

- closed source
- neutron tracking with energy groups
- switches for features like light ion tracking

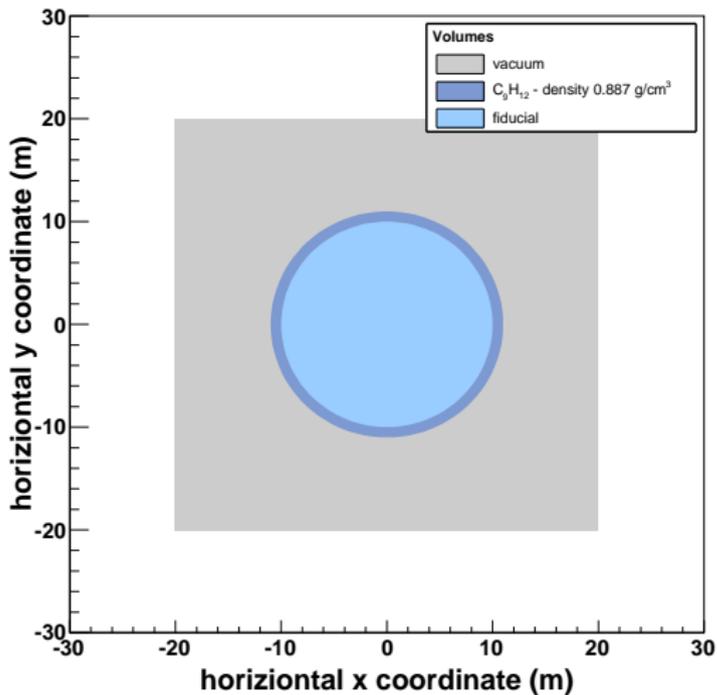
simulating cosmogenics

- from discussion at the previous AARM workshop a series of simulations were suggested with simple geometry but varied materials and μ energies
- with $\mu^{+/-}$ incident on a cylinder radius 11m and thickness 3200 g/cm² with axis along \hat{z}
- neutron counting confined to inner 10m radius and 1600 g/cm² to select an equilibrated volume

simulating cosmogenics (2)



simulating cosmogenics (3)



materials

Material Name	Chemical Composition	Density (g/cm^3)
Liquid Scintillator	C_9H_{12}	0.887
Water	H_2O	1.0
Salt	NaCl	2.17
Carbon	C	2.267
Calcium Carbonate (calcite)	CaCO_3	2.71
Iron	Fe	7.874
Lead	Pb	11.342

Table: Table of materials and densities

statistics

Material Name	Energy (GeV)	Geant4 Primaries (μ^-/μ^+)	FLUKA Primaries (μ^-/μ^+)
Liquid Scintillator	10	600k/600k	1.2M/-
Liquid Scintillator	30	100k/100k	1.2M/-
Liquid Scintillator	100	200k/200k	1.2M/-
Liquid Scintillator	280	1.2M/1.3M	1.2M/-
Liquid Scintillator	1000	10k/-	1.2M/-
Lead	10	550k/550k	800k/-
Lead	30	550k/550k	1.1M/-
Lead	100	530k/510k	1.2M/-
Lead	280	530k/510k	500k/-
Lead	1000	441k/222k	-/-

Table: Table of simulation statistics

our goals

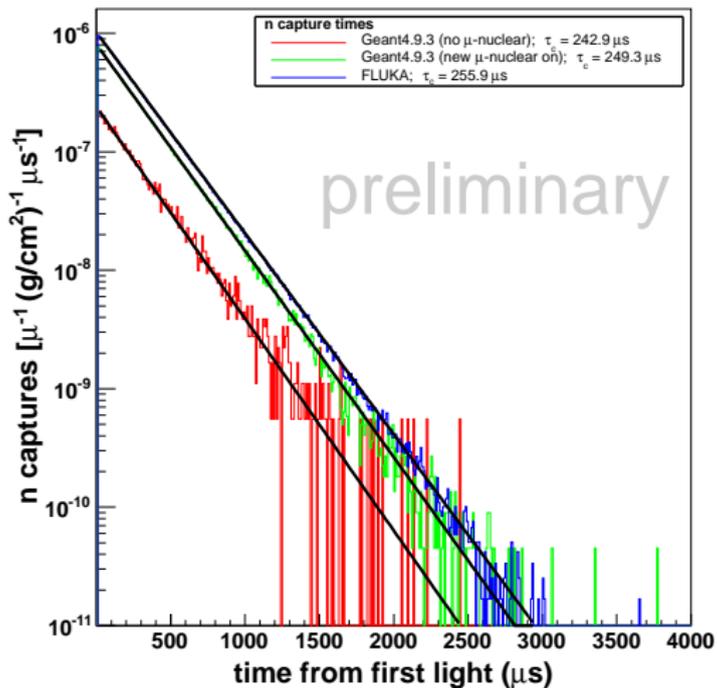
AARM Feb. 2011 charge

- count neutrons in materials without double counting
- obtain information about reaction mechanisms, μ spallation, π spallation ...
- simulations at rock/cavern boundaries with empty cavern

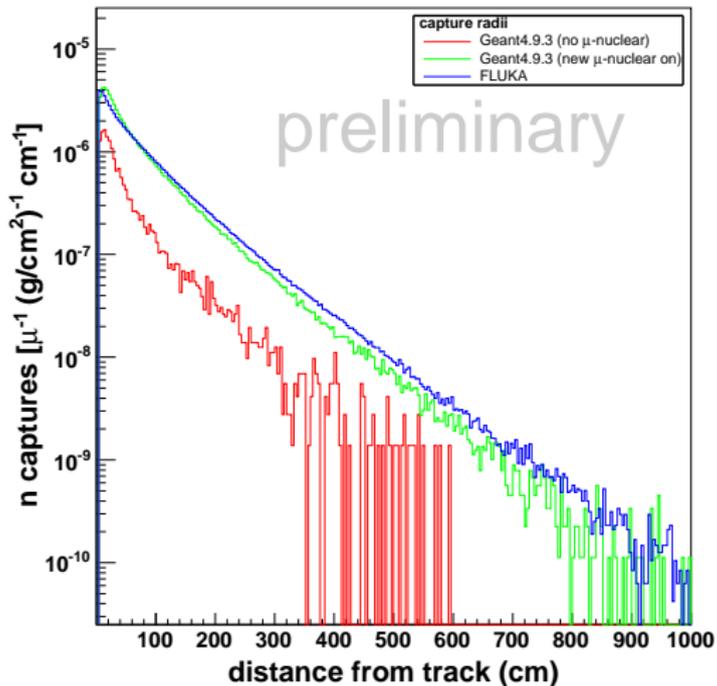
further goals

- find the specific reasons for differences in Geant4 and FLUKA
- understand the microscopic processes and how they are handled in both codes for large cascades
- through these excersizes find the most efficient way to make use of various data that has been taken at underground sites

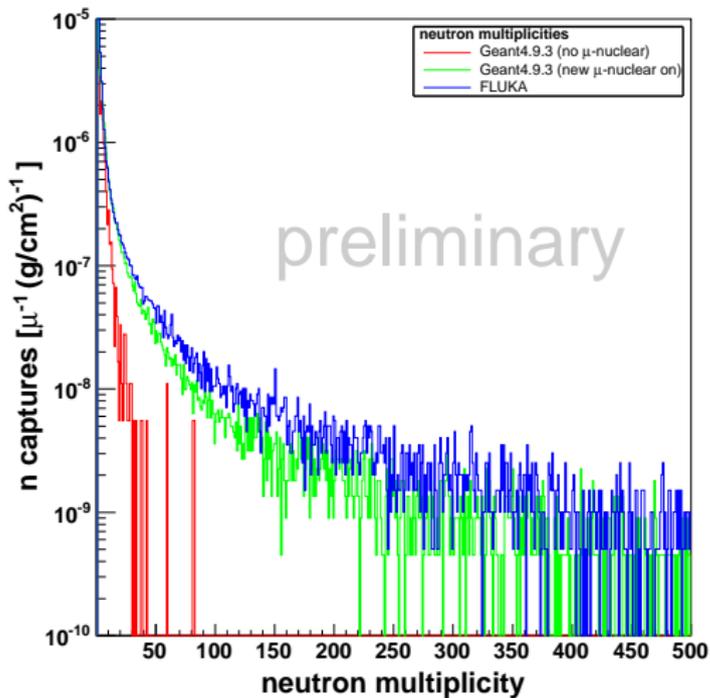
scintillator with 280 GeV μ^-



scintillator with 280 GeV μ^- (2)



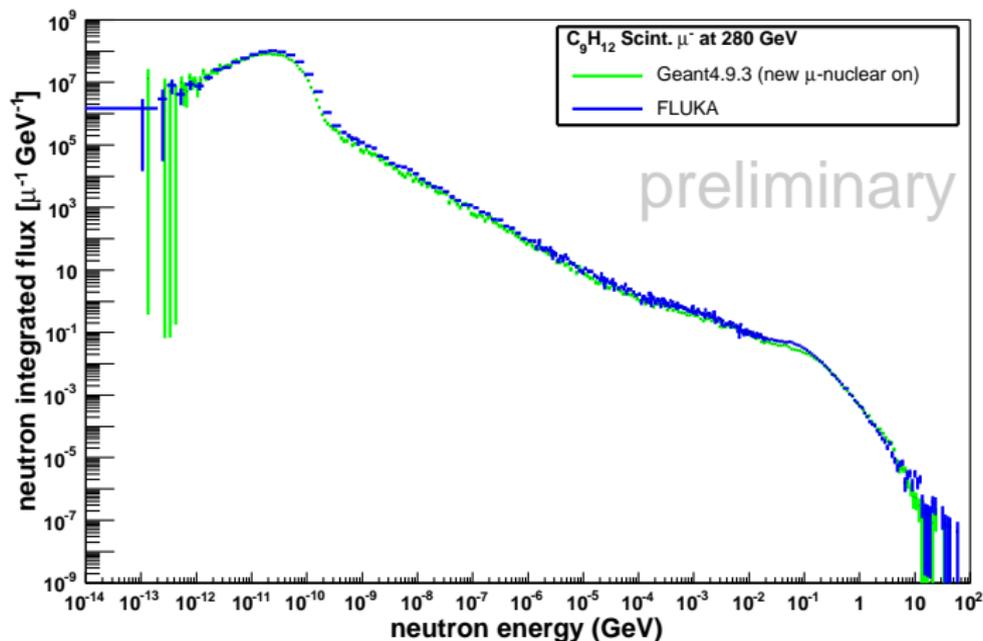
scintillator with 280 GeV μ^- (3)



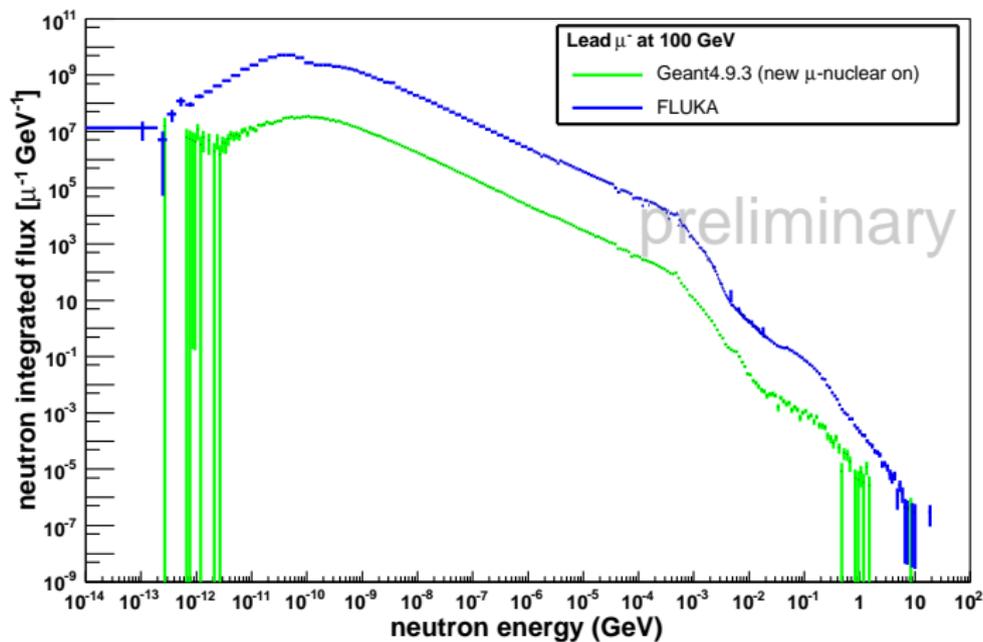
scintillation results

- it is clear that the μ -nuclear interaction is very important to produce accurate neutron spectra, especially for capture multiplicity
- the CDMS collaboration found that the μ -nuclear interaction included in Geant v4.9.3 *was not* producing enough neutrons but the new version of the interaction comes closer
- with the Geant v4.9.3+ we still have under-prediction of neutrons from capture
- there might be a slight mismatch in neutron capture time between Geant v4.9.3 and FLUKA
- simulation data will be used to isolate simulation differences

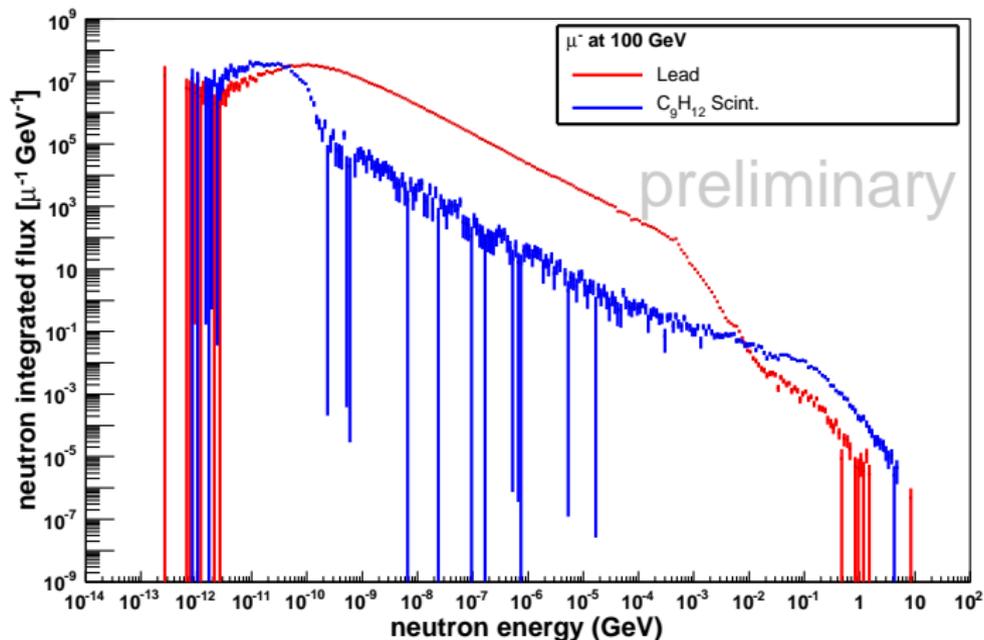
energy flux comparison



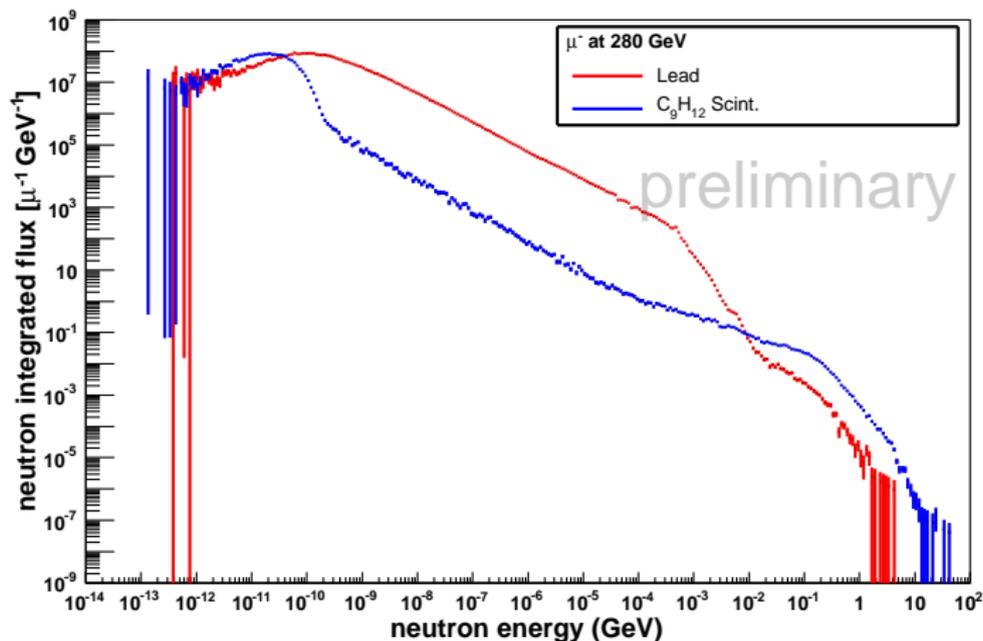
energy flux comparison (2)



energy flux comparison (3)



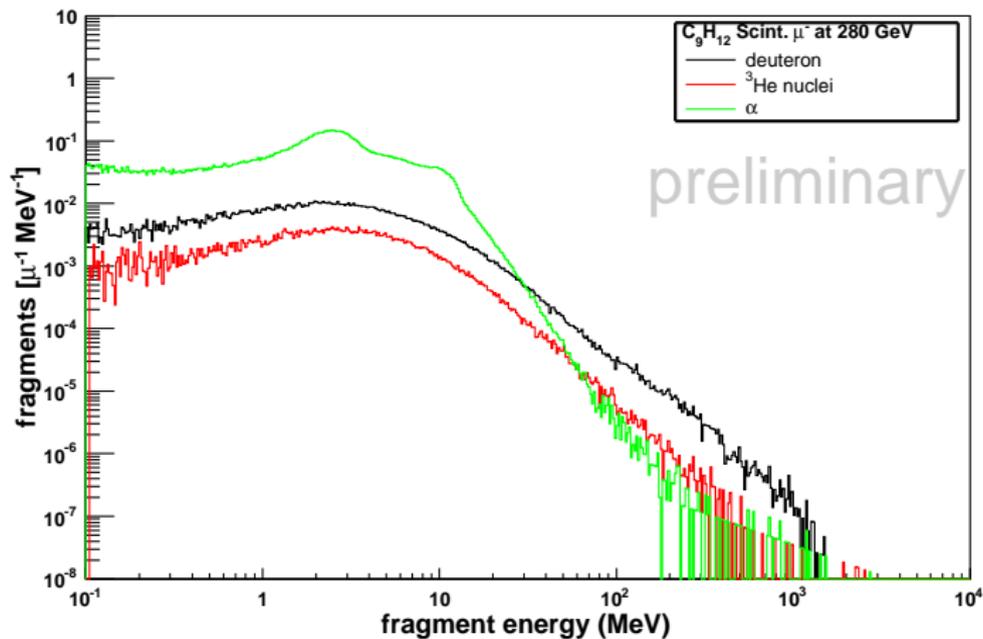
energy flux comparison (4)



energy flux results

- for 280 GeV scintillator the energy flux agrees quite well with some minor Geant v4.9.3+ under-prediction in the 1 GeV neutron energy range
- for 100 GeV lead the energy flux disagrees by a large factor
- it is clear that at the high energy for Geant v4.9.3+ the scintillator and lead spectra do not overlap and scintillator actually out-produces lead
- these results are **preliminary** and will be scrutinized in detail before any judgment about either FLUKA or Geant is made

fragment analysis with 280 GeV μ^-



fragment analysis results

- light fragments are not produced by default in FLUKA and it is not clear that the deuteron is treated correctly if it is produced
- since deuterons are very loosely bound when production of neutrons is concerned high energy deuterons are also of concern
- preliminary investigations show that according to Geant deuterons can be more abundant than the He isotopes at high energies even when the primary nucleus is an α -cluster nucleus (Carbon).

summary

- Geant (v4.9.3+) and FLUKA (v2011.2.5) are used to produce cosmogenic simulations in a cylindrical geometry
- materials and energies are varied in these simulations **lots of statistics coming in with SLAC batch farm!**
- it is clear that Geant v4.9.3 was under-producing compared to FLUKA and the new μ -nuclear interaction addresses that to a large extent
- this new interaction is in the “special” Geant v4.9.3+ but not yet in general release (v4.9.5)
- going forward to produce large statistics with above versions and all materials
- more detailed flux/production mechanism analysis expected end of this year