1	Radiogenic Neutron Dackground
2	v2.0
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11	Ultra-low-background experiments, generically termed rare-event searches, address some of the
12	most important open questions in particle physics, cosmology and astrophysics: direct detection of
13	dark matter, neutrinoless double beta decay, proton decay, and detection of solar and supernovae
14	neutrinos. Although their detection methods and physics goals are varied, rare-event searches share a
15	number of common requirements in order to obtain sensitivity to low-rate processes in the presence of
16	an overwhelming rate of environmental radiation, including the need for significant rock overburdens
17	to moderate the flux of cosmic-ray muons. This requirement is so generic that an international
18	community of underground science has emerged, with dozens of experiments operating in mines
19	that vary in depth from a few hundred meters water equivalent (m.w.e.) to many thousand.
20	Simulations are used to understand backgrounds caused by naturally occurring radioactivity in the
21	rock and in every piece of shielding and detector material used in the experiment. Most important
22	are processes like spontaneous fission and (α, n) reactions in material close to the detectors that can

and spectra obtained with Mei-Zhang-Hime and SOURCES4 codes are presented.

Badiogenic Neutron Background

produce neutrons. A comparison study between two dedicated softwares is detailed. Neutron yields

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I. INTRODUCTION

Reduction of radiogenic backgrounds is one of the most 26 important factors for rare event search experiments in-27 cluding searches for dark matter and neutrino-less dou-28 ble beta decay. These radiogenic backgrounds can be 29 generically classified into two types: electron recoil and 30 nuclear recoil backgrounds. Electron recoil background 31 result from the interactions of gamma rays, electrons or 32 beta particles interacting with the electrons in the detec-33 tor's target medium while nuclear recoil backgrounds re-34 sult from neutrons interacting with the nucleus in the de-35 tector's target medium. In dark matter experiments, the 36 nuclear recoil background is of particular concern because 37 a nuclear recoil from a neutron can be indistinguishable 38 from a nuclear recoil from a WIMP. 30

Neutrons are produced by spontaneous fission, α -40 n interactions and muon-induced interactions. Muon-41 induced neutrons can be reduced by operation the de-42 tector deep underground and by placing passive shielding 43 ⁴⁴ around the detector. In addition, muon-induced neutrons ⁴⁵ can often be recognized by identifying the parent muon in a muon veto. More difficult to deal with are neutrons 46 resulting from α -n interactions and spontaneous fission 47 ⁴⁸ from ²³⁸U, ²³⁵U, and ²³²Th present in the materials the shielding and the detectors themselves. Thus, experi-49 50 mentalists must select their materials carefully when de-51 signing and constructing their experiments. This requires that they have accurate simulations of the processes in 52 53 the materials they are considering.

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⁵⁶ by Mei, Zhang and Hei which we will refer to as the USD 57 code.

II. NEUTRON YIELD AND SPECTRA

59 The measurements of neutron spectra strongly depend 60 on the material and are not straightforward since neu-⁶¹ trons are neutral particles. Their calculations is criti-62 cal to low background experiments such as direct dark 63 matter search experiments. The total neutron yield in-⁶⁴ dicates the number of neutrons which are produced or ⁶⁵ had entered the target whereas the neutron energy spec-⁶⁶ trum determines the background events we would expect ⁶⁷ in the energy range of interest. Both are then needed 68 in order to carry out a complete and reliable neutron ⁶⁹ background simulation. The evaluation of neutron yields 70 and spectra can be performed via different codes: the 71 modified version of SOURCES4A [1] and the code de-⁷² veloped by Mei, Zhang and Hime [2] made available 73 online at http://neutronyield.usd.edu have been consid-74 ered. SOURCES4 code has been modified to extend the ⁷⁵ cross section for (α, n) up to 10 MeV, based on experimen-⁷⁶ tal data, whenever possible, and calculations performed ⁷⁷ via the EMPIRE code [3] by the group at the Univer-78 sity of Sheffield. SOURCES4A and USD codes calculate ⁷⁹ neutron yields and spectra from (α, n) reactions due to ⁸⁰ the decay of radionuclides. Radiogenic neutrons are due ⁸¹ to intrinsic contamination of materials surrounding the ⁸² detectors of ²³²Th, ²³⁸U and ²³⁵U decay chains. Compar-There are currently two codes available for use in such ⁸³ ison study between SOURCES4 and USD codes is car-⁵⁵ calculations, SOURCES-4 and a second code developed ⁸⁴ ried out considering ²³²Th, ²³⁸U decay chains in secular

85 equilibrium, although a possibility of disequilibrium can ⁸⁶ be taken into account: due to migration differently, the 87 long-lived isotopes, ²²⁶Ra, ²²²Rn, ²¹⁰Po, ²²⁸Ra, ²²⁸Th ⁸⁸ and the associated decay daughters could be calculated ⁸⁹ separately, see table I. For both chains, most of the neu-⁹⁰ trons are produced by the α generated in the second part 91 of the chains.

TABLE I. Radiogenic neutron yield $(n \cdot s^{-1} \cdot cm^{-3})$ from (α, n) reactions in different materials for 1ppb of ²³⁸U and ²³²Th decay chains. Neutron yield have been calculated via the modified SOURCES4A code. -3)

Neutron yield for 1ppb (n·s ⁻¹ ·cn				
Material	$^{238}\mathrm{U} ightarrow ^{226}\mathrm{Ra}$	226 Ra $\rightarrow ^{206}$ Pb		
Stainless Steel	$6.4 \cdot 10^{-15}$	$3.1 \cdot 10^{-11}$		
Pyrex	$4.0 \cdot 10^{-11}$	$1.9 \cdot 10^{-10}$		
Borosilicate Glass	$6.3 \cdot 10^{-11}$	$2.8 \cdot 10^{-10}$		
Titanium	$1.14 \cdot 10^{-13}$	$1.0 \cdot 10^{-10}$		
Copper	$0.0 \cdot 10^{-11}$	$2.8 \cdot 10^{-12}$		
$PE(C_2H_4)$	$1.6 \cdot 10^{-12}$	$1.1 \cdot 10^{-11}$		
PTFE (CF_2)	$1.8 \cdot 10^{-10}$	$1.6 \cdot 10^{-9}$		
	$^{232}\mathrm{Th} ightarrow ^{228}\mathrm{Th}$	$^{228}{ m Th} ightarrow ^{208}{ m Pb}$		
Stainless Steel	$8.8 \cdot 10^{-19}$	$4.1 \cdot 10^{-11}$		
Pyrex	$2.4 \cdot 10^{-12}$	$8.4 \cdot 10^{-11}$		
Borosilicate Glass	$3.8 \cdot 10^{-12}$	$1.2 \cdot 10^{-10}$		
Titanium	$4.4 \cdot 10^{-16}$	$9.3 \cdot 10^{-11}$		
Copper	$0.0 \cdot 10^{-11}$	$9.5 \cdot 10^{-12}$		
$PE(C_2H_4)$	$1.6 \cdot 10^{-13}$	$5.1 \cdot 10^{-12}$		
PTFE (CF_2)	$7.1 \cdot 10^{-12}$	$7.7 \cdot 10^{-10}$		

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The calculation of the neutron spectra requires as in-94 put the cross-sections of (α, n) reactions, the probabilities 95 of nuclear transition to different excited states (branch-96 ing ratio) and the alpha emission lines from the radioac-97 tive radionuclides. Both codes consider a thick target: 98 calculation of neutron yields and spectra are carried out ۵Q under the assumption that the size of radioactive sample exceeds significantly the range of alpha. The energy bin 101 size of the (α, n) calculation is fixed to be 0.1 MeV in 102 USD code while is user dependent in SOURCES4A. 103

Table II lists α -lines respectively for ²³²Th and ²³⁸U 104 decay chains, present in SOURCES4A and USD codes. 105 USD code is missing the lines coming from ²²²Ra iso-106 tope in ²³⁸U decay chain which SOURCES4A library is 107 considering. Overall, the branching ratio and the energy 108 lines are in good agreement between the two codes. 109

Cross-sections and branching ratios are required as well $_{122}$ perimentally observed. 110 for neutron yield and spectra calculations. The USD 123 111 ¹¹² code considers TENDL 2012 [4] to provide (α , n) nu-¹²⁴ section libraries used in USD and SOURCES4A has been 113 114 115 116 ¹¹⁷ and, for some isotopes, a combination of data measure- ¹²⁹ such as stainless steel, copper, titanium, borosilicate glass ¹¹⁸ ments and EMPIRE2.19 calculations. Also, EMPIRE is ¹³⁰ (PMTs glass), PTFE which become important when the ¹¹⁹ the code recommended by International Atomic Energy ¹³² external flux is attenuated by the shielding. Compari-¹²⁰ Agency (IAEA). Neither EMPIRE nor TALYS can cal-¹³³ son of cross section inputs results in a good agreement ¹²¹ culate properly resonance behavior which has been ex- ¹³⁴ for most of isotopes in both code libraries. For some,

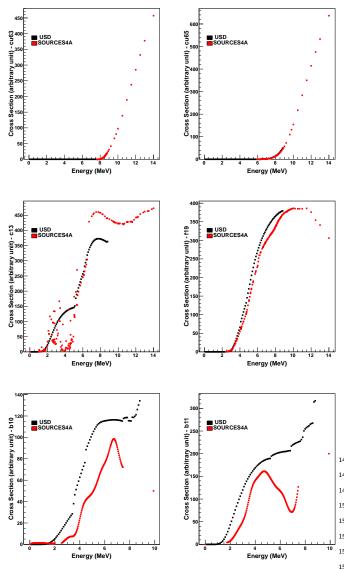
TABLE II. Alpha lines present in SOURCES4A and USD, and their intensity (BR) for isotopes in the 232 Th and 238 U decay chains. Only lines with intensity > 1% have been quoted.

	SOURC	ES4A	USD			
Isotopes	Line (keV)	BR (%)	Line (MeV)	BR (%)		
238 U	4151	21	4.151	21		
U	4198	79	4.198	79		
234 U	4722.4	28.42	4.722	28.6		
U	4774.6	71.38	4.775	71.4		
230 Th	4620.5	23.4	4.621	23.7		
111	4687.0	76.3	4.688	76.3		
226 Ra	4601	5.55	4.602	5.6		
	4784.34	94.45	4.784	94.4		
222 Rn	5489	99.92	5.490	100		
218 Po	6002.35	99.98	6.002	100		
214 Po	7686.82	99.99	7.687	99.99		
$^{210}\mathbf{Po}$	5304.33	99.99	5.304	100		
232 Th	3947.2	21.7	3.954	22.1		
111	4012.3	78.2	4.013	77.9		
228 Th	5340.36	27.2	5.340	28.5		
111	5423.15	72.2	5.423	71.5		
224 Ra	5448.6	5.06	5.449	5.1		
	5685.37	94.92	5.685	94.9		
220 Rn	6288.3	99.99	6.288	100		
$^{216}\mathbf{Po}$	6778.5	99.99	6.778	100		
212 Bi	6051.1	25.16	6.050	26.2		
	6090.2	9.79	6.0902	9.8		
212 Po	8784.6	100	8.784	64		

TABLE III. Radiogenic neutron yield $\rm (n/s/cm^3)$ for copper and polyethylene materials and for $^{238}\rm U$ and $^{232}\rm Th$ decay chains. Column (1) and (2) refer to pure USD and SOURCES4A calculation, respectively. Column (3) refers to SOURCES4A calculation with USD (α , n) cross section libraries. A ratio of the neutron yield is also provided: column (a) refers to the ratio of (2) over (1), whereas column (b)corresponds to the ratio (2)/(3)

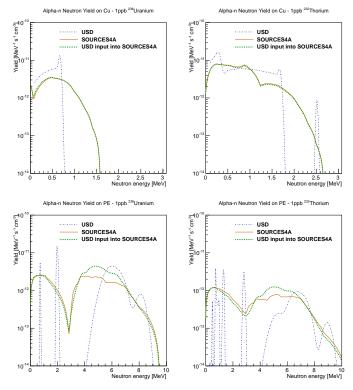
	$\begin{array}{c} \textbf{Neutron Yield} \\ \textbf{(10}^{-12} \cdot \mathbf{n} \cdot \mathbf{s}^{-1} \cdot \mathbf{cm}^{-3} \textbf{)} \end{array}$			Ratio		
Material	Chain	(1)	(2)	(3)	(a)	(b)
	^{238}U ^{232}Th	$3.46 \\ 11.1$	$2.84 \\ 9.49$	$2.93 \\ 9.18$	$\begin{array}{c} 0.8 \\ 0.9 \end{array}$	$\begin{array}{c} 1.0\\ 1.0\end{array}$
Polyethylene (C_2H_4)	²³⁸ U ²³² Th	$9.56 \\ 2.87$	$12.6 \\ 5.28$	$16.4 \\ 5.97$	$1.3 \\ 1.8$	$\begin{array}{c} 0.8 \\ 0.9 \end{array}$

An extensive comparison between the different crossclear cross-sections. TENDL is a validated nuclear data 125 carried out for the target nuclides present in the matelibrary which provides the output of the TALYS [5] nu- 126 rials contributing the most to radiogenic neutron backclear model code system; SOURCES4 cross section in- 127 ground. Specifically, we refer to materials which comput libraries come from EMPIRE2.19 [3] calculations ¹²⁸ pose the shielding scheme and the internal detector parts,



 $_{136}$ crepancies. To better understand the contribution due $_{158}$ SOURCES4A (column (2)) input (α , n) cross-sections. 137 to the cross section input library in the radiogenic neu- 159 For polyethylene, we can conclude that the input cross-¹³⁸ tron yield and spectrum we have calculated radiogenic ¹⁶⁰ section may account up to a 20% discrepancy in neutron ¹³⁹ neutron yield and spectra for two different materials: ¹⁶¹ yield. Figure 3 shows radiogenic neutron spectra for cop-¹⁴⁰ copper, for which the input cross sections in both codes ¹⁶² per (upper row) and polyethylene (lower row), both from 141 are matching, and polyethylene, for which input cross 163 uranium and thorium decay chains, left and right panels ¹⁴² sections of ¹³C show discrepancies between the codes, ¹⁶⁴ respectively. ¹⁴³ details in figure 1. We have considered natural copper ¹⁶⁵ 144 (70% ⁶³Cu and 30% ⁶⁵Cu) with a density of 8.96 g/cm³; 166 obtained via SOURCES4A and USD codes for different $_{145}$ polyethylene material (C₂H₄) is considered with a density $_{167}$ material has been carried out. Results are shown in ta-¹⁴⁶ of density is 0.935 g/cm³. Estimates are done for 1ppb ¹⁶⁸ ble IV and figure 3. A qualitative agreement between

FIG. 2. Radiogenic neutron spectra $(n \cdot MeV^{-1} \cdot s^{-1} \cdot cm^{-3})$ calculated for 1ppb ²³⁸U and ²³²Th decay chains, left and right panels, respectively. First row show copper contribution, lower row polyethylene material. Dotted blu lines refers to pure USD calculations, plain orange line to pure SOURCES4A calculations and dashed green line is the mixed computation for which we ran SOURCES4A algorithm with input cross section of USD code.



147 in ²³⁸U and ²³²Th decay chains. Calculations consist of ¹⁴⁸ pure SOURCES4A and USD computation and a mixed 149 computation for which we run SOURCES4A algorithm ¹⁵⁰ with cross section inputs of USD. Resulting radiogenic ¹⁵¹ neutron yields are quoted in table III. The column (a) of ¹⁵² the radiogenic neutron yield ratio quotes the calculation ¹⁵³ differences between SOURCES4A and USD codes: they ¹⁵⁴ show a reasonably good agreement, inside a 50% discrep-¹⁵⁵ ancy. The column (b) refers to the ratio of radiogenic ¹⁵⁶ neutron yield resulting from the same algorithm calcu-¹³⁵ such as ¹³C and ¹⁰B and ¹¹B the cross sections show dis-¹⁵⁷ lation (SOURCES4A) but with USD (column (3)) and

A comparison of neutron yield and energy spectrum

¹⁷⁰ a factor 2 is found. The energy spectra calculated via ¹⁷⁷ framework with Geant4.9.5.p01 and the pertinent high 171 SOURCES4A code are in general smoother, without the 178 precision neutron physics list utilizing cross sections from ¹⁷² presence of peaks, feature of USD spectra.

TABLE IV. Radiogenic neutron yield $(n \cdot s^{-1} \cdot cm^{-3})$ per material considering 1ppb of ²³⁸U and ²³²Th. The percentage difference is calculates as (SOURCES4A-USD)/[(SOURCES4A+USD)/2].

	$\frac{\textbf{Neutron Yield}}{(\textbf{n} \cdot \textbf{s}^{-1} \cdot \textbf{cm}^{-3})}$			
Material	Chain	SOURCES4	A USD	Diff %
Cu	$^{238}\mathrm{U}$	2.84E-12	3.46E-12	20
Cu	$^{232}\mathrm{Th}$	9.49E-12	1.11E-11	16
$PE(CH_2)$	^{238}U	1.26E-11	9.56E-12	-27
$1 \ge (0.112)$	232 Th	5.28E-12	2.87E-12	-59
Titanium	²³⁸ U	$1.04 \ 10^{-10}$	$1.99 \ 10^{-10}$	-63
1 Itamum	$^{232}\mathrm{Th}$	$9.29 \ 10^{-11}$	$1.24 \ 10^{-10}$	-28
Stainless Steel	^{238}U	$3.10 \ 10^{-11}$	$5.95 \ 10^{-11}$	-63
Stamless Steel	$^{232}\mathrm{Th}$	$4.05 \ 10^{-11}$	$6.80 \ 10^{-11}$	-51
Pyrex	^{238}U	$2.30 \ 10^{-10}$	$1.61 \ 10^{-10}$	36
I ylex	232 Th	$8.66 \ 10^{-11}$	$4.59 10^{-11}$	61
Borosilicate Glass	^{238}U	$3.48 \ 10^{-10}$	$2.45 \ 10^{-10}$	35
Dorosilicate Glass	$^{232}\mathrm{Th}$	$1.27 10^{-10}$	$6.98 \ 10^{-11}$	58
PTFE (CF_2)	^{238}U	$1.81 \ 10^{-9}$	$1.60 \ 10^{-9}$	12
$\Gamma \Gamma \Gamma \Gamma (O \Gamma 2)$	$^{232}\mathrm{Th}$	$7.76 \ 10^{-10}$	$5.42 \ 10^{-10}$	36

III. **GEANT4 PROPAGATION**

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174 To evaluate the impact of the varying neuron spectra ¹⁷⁵ produced by SOURCES4A and the USD calculator, sim-

169 the two codes is observed. A maximum discrepancy of 176 plified detector geometries were created within a RAT 179 G4NDL3.14 for neutrons under 20 MeV. Four simula-180 tions, for neutrons from uranium and throrium, with the ¹⁸¹ spectra from SOURCES4A and USD (shown in Figure 182 3.) were run for each relevant material under study.

> Three simplified geometries were created to study neu-183 ¹⁸⁴ tron elastic scatter signals within central detector mate-185 rials and the potential for vetoing neutrons with outer 186 detectors. A spherical liquid argon detector, with a radius of 1 m, is surrounded by shells of 10 cm thick acrylic, 187 5 mm thick borosilicate glass, and water out to a radius 189 of 3 m. The simulated neutrons were isotropically cre-¹⁹⁰ ated in the borosilicate glass, as it is the leading source ¹⁹¹ of radiogenic neutrons in many liquid argon detectors.

> The second geometry studied was a cylindrical liquid 192 xenon detector with a 1 m diameter and height. It is 193 194 nested within cylinders of 3 cm thick PTFE, 2 cm thick 195 titanium, and liquid scintillator with a diamteriater and height of 3 m. Separte neutron simulations were done for spectra from titanium and PTFE. plots need to be ¹⁹⁸ added, references need to be added

IV. CONCLUSION

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FIG. 3. Radiogenic neutron spectra ($n \cdot MeV^{-1} \cdot s^{-1} \cdot cm^{-3}$) calculated for 1ppb ²³⁸U and ²³²Th decay chains, left and right panels, respectively. The (α , n) reaction contribution is shown in various commonly used materials from SOURCES4A in orange and USD in blue. From top to bottom materials are: copper, titanium, stainless steel, pyrex, borosilicate glass, polyethylene and teflon (PTFE).

